

MIMAC

Micro-tpc MAtrix of Chambers

A Large TPC for Directional Dark Matter detection

Daniel Santos

Laboratoire de Physique Subatomique et de Cosmologie

(LPSC-Grenoble)

(Université Grenoble-Alpes -CNRS/IN2P3)



MIMAC (Micro-tpc MAtrix of Chambers)

LPSC (Grenoble) : D. Santos, F.Naraghi , N. Sauzet (CDD)

-Technical Coordination, Gas circulation and detectors : **O. Guillaudin**

- Electronics : **G. Bosson, J. Bouvier, J.L. Bouly,**

L.Gallin-Martel, F. Rarbi

- Data Acquisition: **T. Descombes**

- Mechanical Structure : **J. Giraud**

- COMIMAC (quenching) : **J-F. Muraz**

IRFU (Saclay): P. Colas, I. Giomataris

CCPM (Marseille): C. Tao, J. Busto

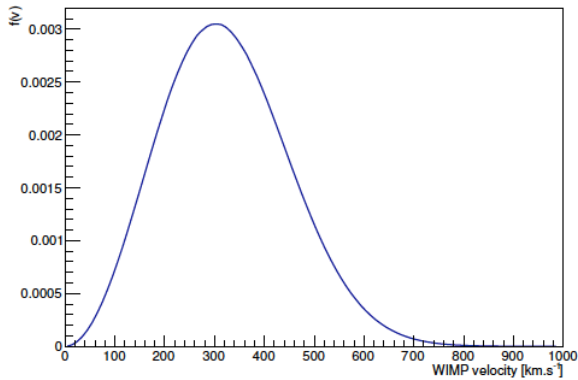
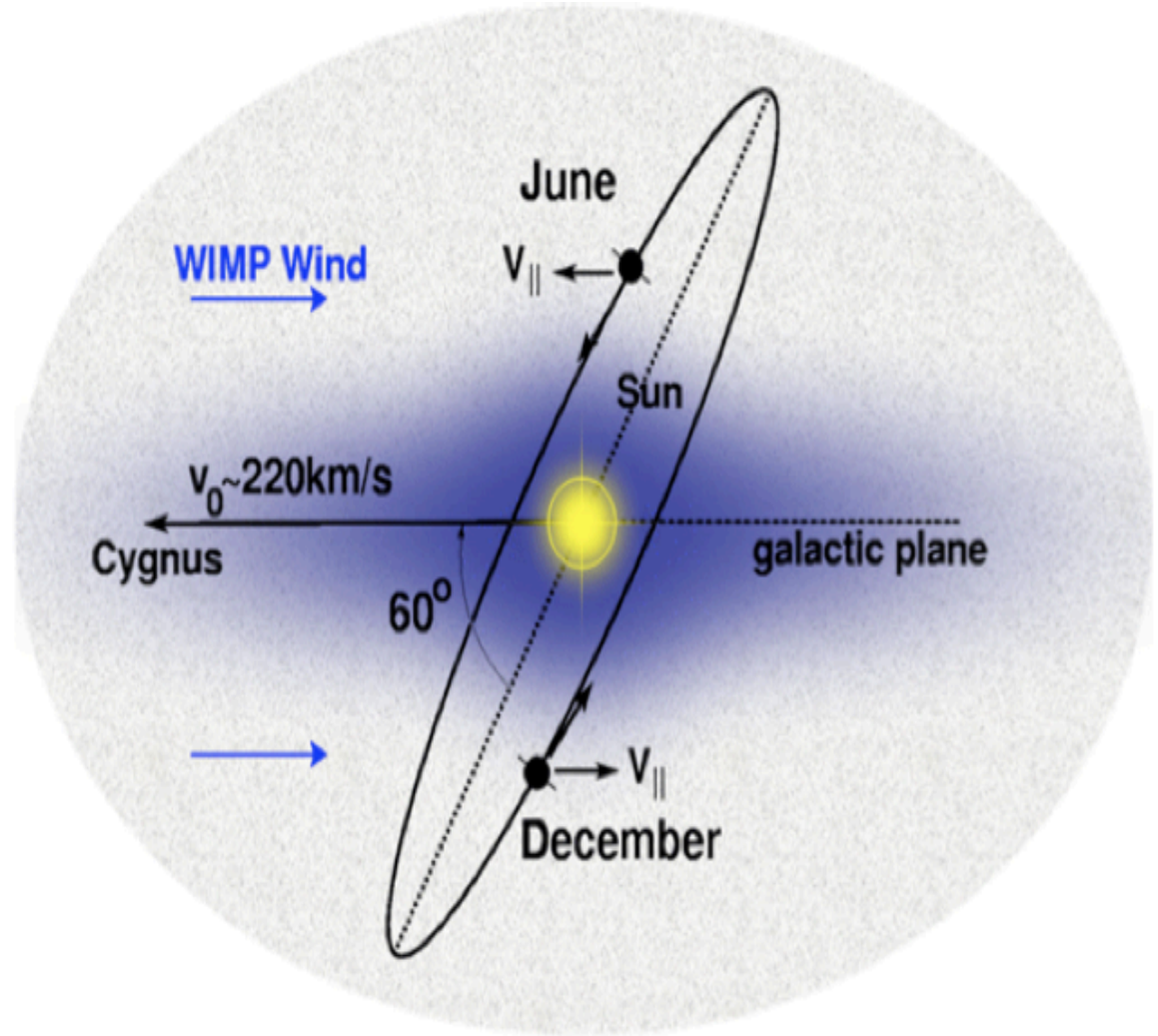
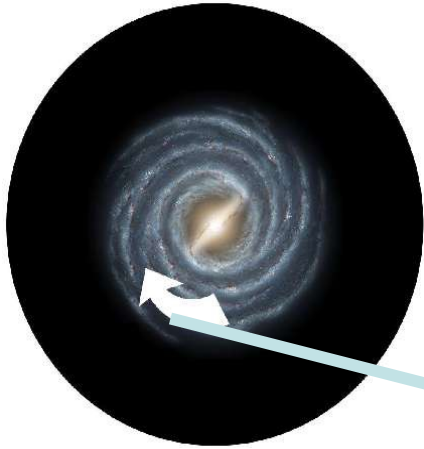
Tsinghua University (Beijing-China): C. Tao, I. Moric (post-doc), Y. Tao (Ph.D)

Prototype hosted in **IHEP (Beijing-China): ZhiminWang , Changgen Yang**

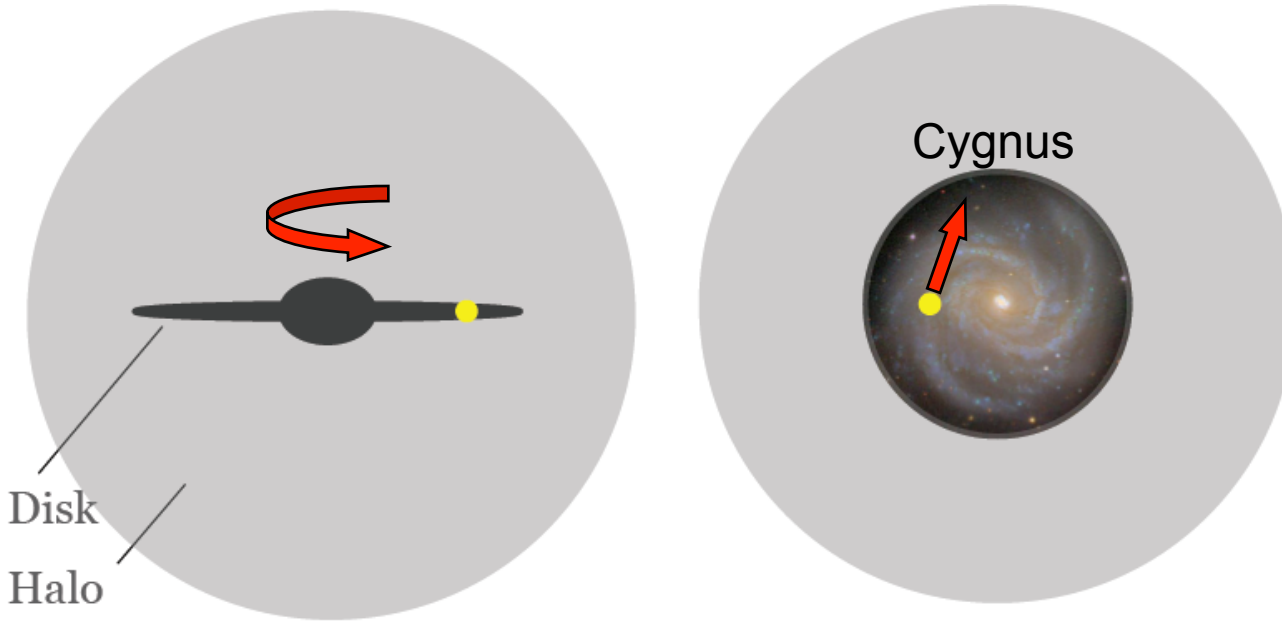
Neutron facility (AMANDE) :

IRSN (Cadarache): V. Lacoste, B. Tampon (Ph. D.)

Directional detection: principle



Directional detection : principle



$$\langle V_{\text{rot}} \rangle \sim 220 \text{ km/s}$$

The signature, the only one (!), able to correlate the events in a detector to the galactic halo !!

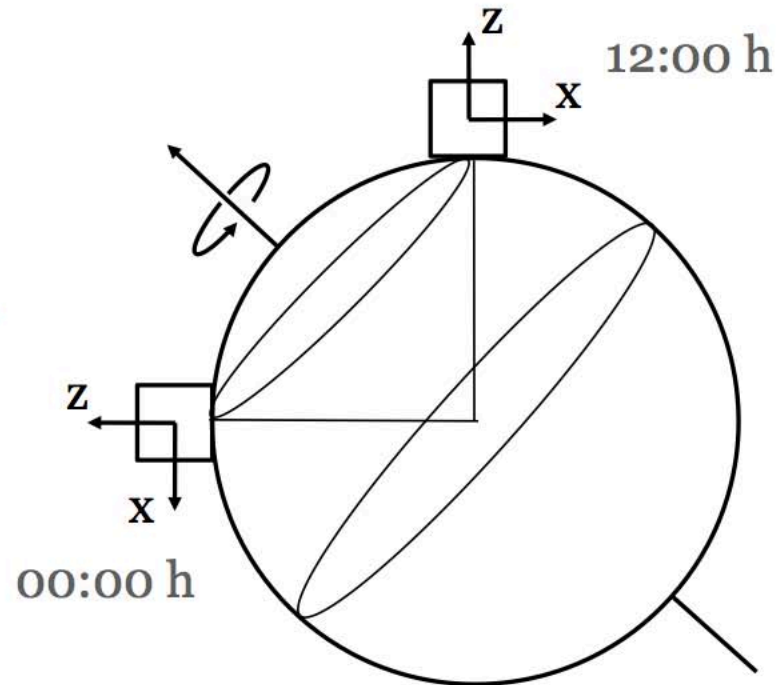
Angular modulation of WIMP flux

Modulation is sidereal (tied to stars) not diurnal (tied to Sun)

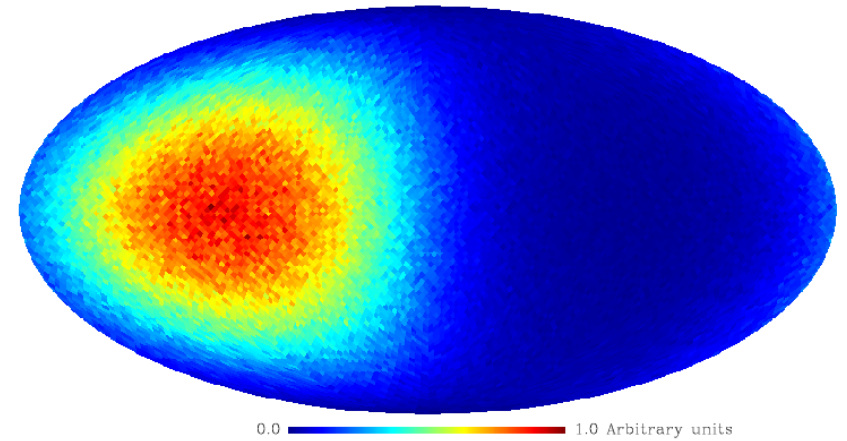
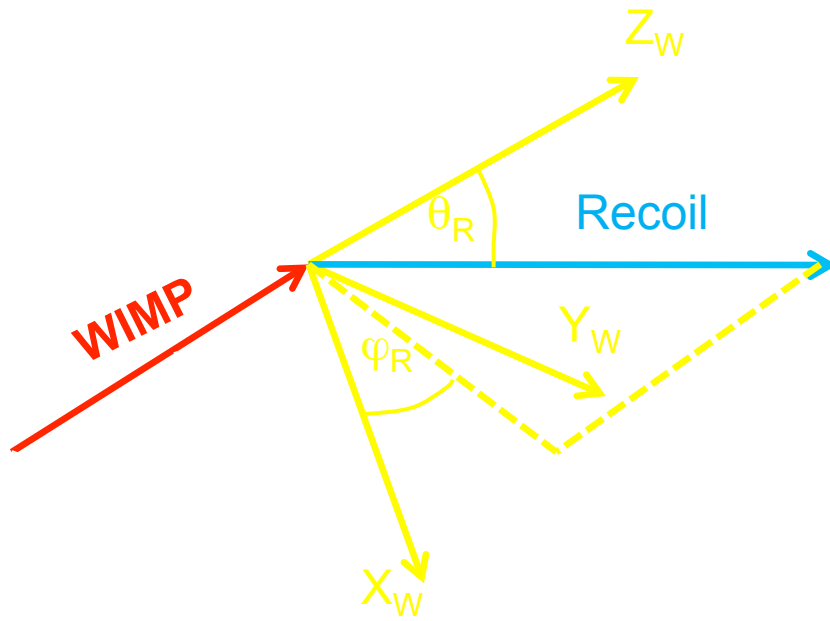
Cygnus



Direction of
Earth motion



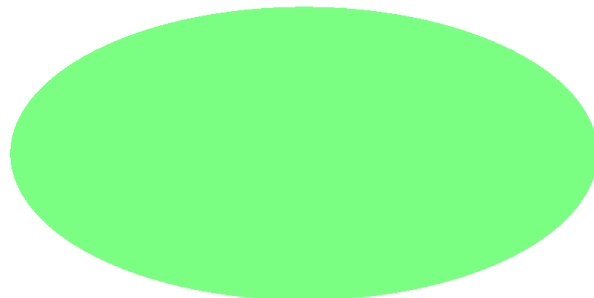
There are many “angles” for nuclear recoils...



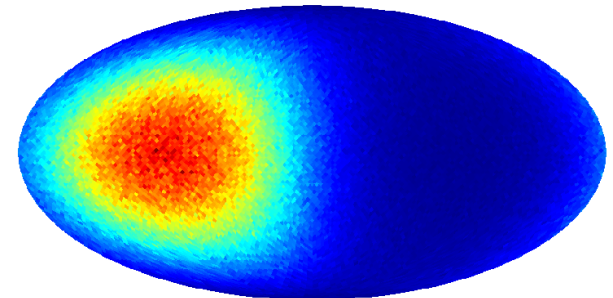
Map of recoils in galactic coordinates (HealPix)

10^8 Events with $E_R = [5, 50]$ keV

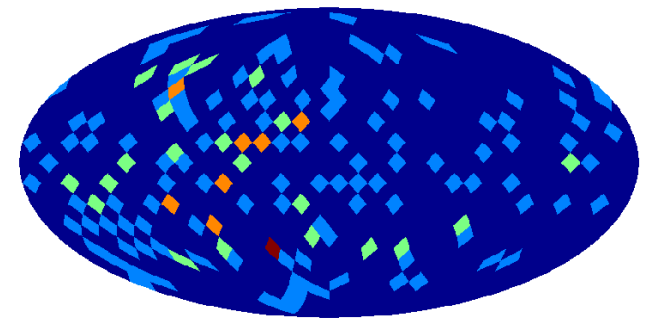
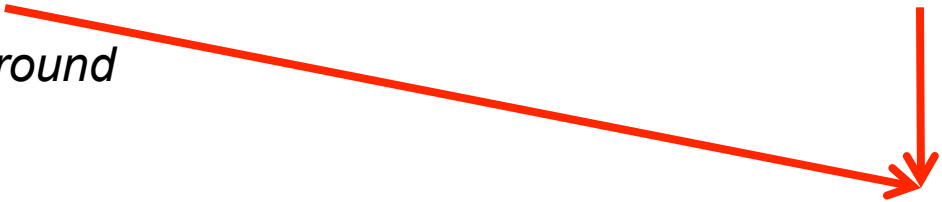
100 WIMP evts + 100 Background evts



Background



Wimp recoils

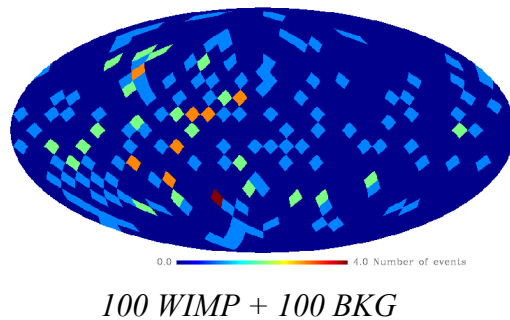


Phenomenology: **Discovery**

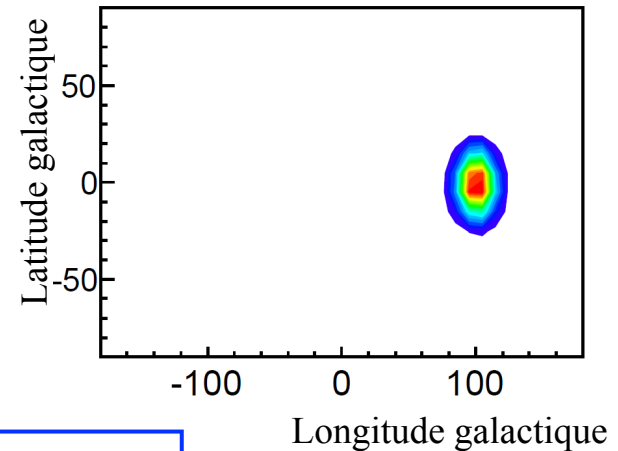
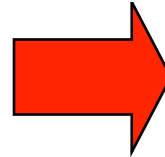
J. Billard *et al.*, PLB 2010
J. Billard *et al.*, arXiv:1110.6079

Proof of discovery: **Signal pointing toward the Cygnus constellation**

Blind likelihood analysis in order to establish the galactic origin of the signal



$$\mathcal{L}(\ell, b, m_\chi, \lambda)$$



Strong correlation with the direction of the Constellation Cygnus even with a large background contamination

Directional Detection : identification

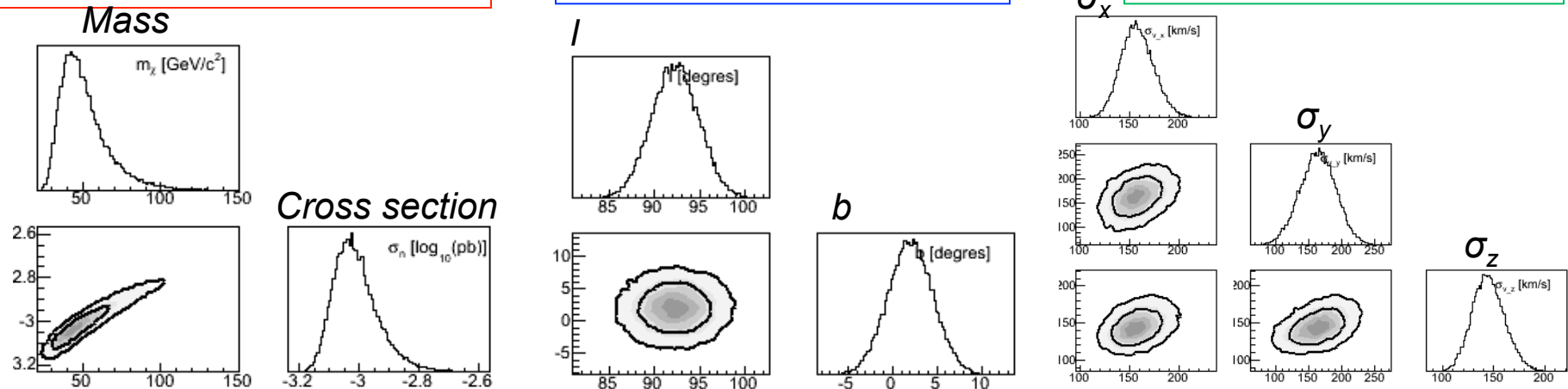
J. Billard *et al.*, PRD 2011

8 parameters simultaneously constrained by only one 3D experiment

Mass – cross section

Dark Matter signature

Galactic Halo shape



	m_χ (GeV/c^2)	$\log_{10}(\sigma_n$ (pb))	l_\odot ($^\circ$)	b_\odot ($^\circ$)	σ_x ($\text{km}\cdot\text{s}^{-1}$)	σ_y ($\text{km}\cdot\text{s}^{-1}$)	σ_z ($\text{km}\cdot\text{s}^{-1}$)	β	R_b ($\text{kg}^{-1}\text{year}^{-1}$)
Input	50	-3	90	0	155	155	155	0	10
Output	$51.8^{+5.6}_{-19.4}$	$-3.01^{+0.05}_{-0.08}$	$92.2^{+2.5}_{-2.5}$	$2.0^{+2.5}_{-2.5}$	158^{+15}_{-17}	164^{+27}_{-26}	145^{+14}_{-17}	$-0.073^{+0.29}_{-0.18}$	10.97 ± 1.2

Directional experiments around the world

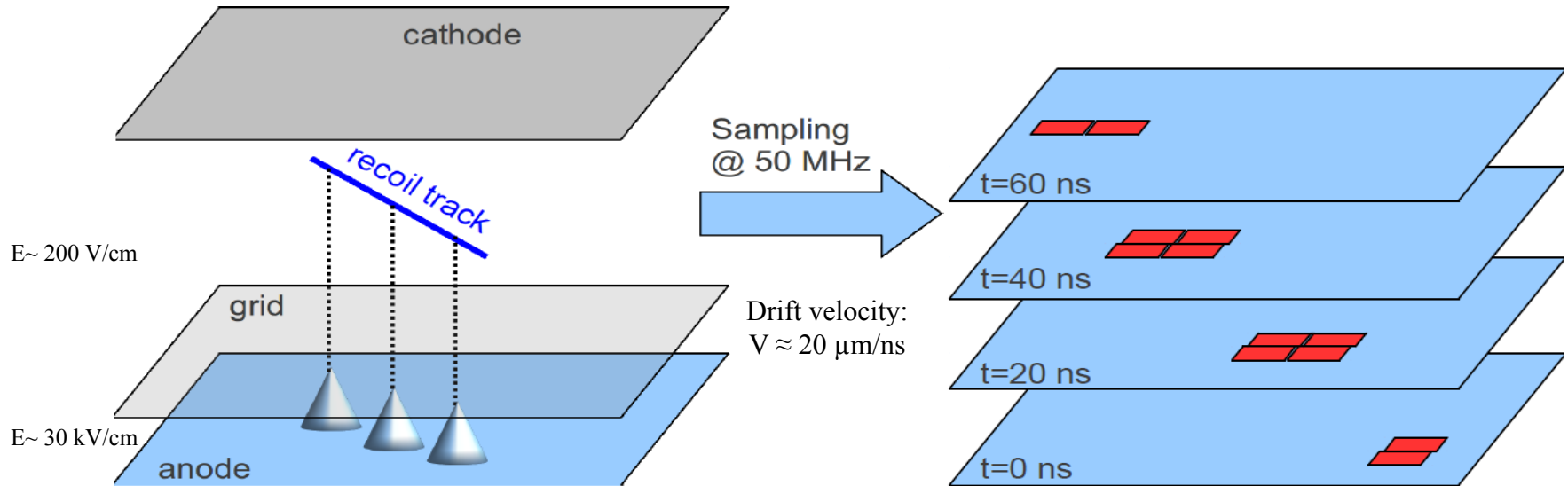


CYGNUS 2017- International Workshop

Xichang, Sichuan (CHINA) – June 13th- 15th 2017



MIMAC: Detection strategy



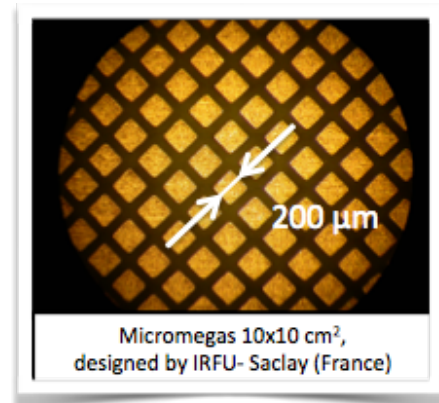
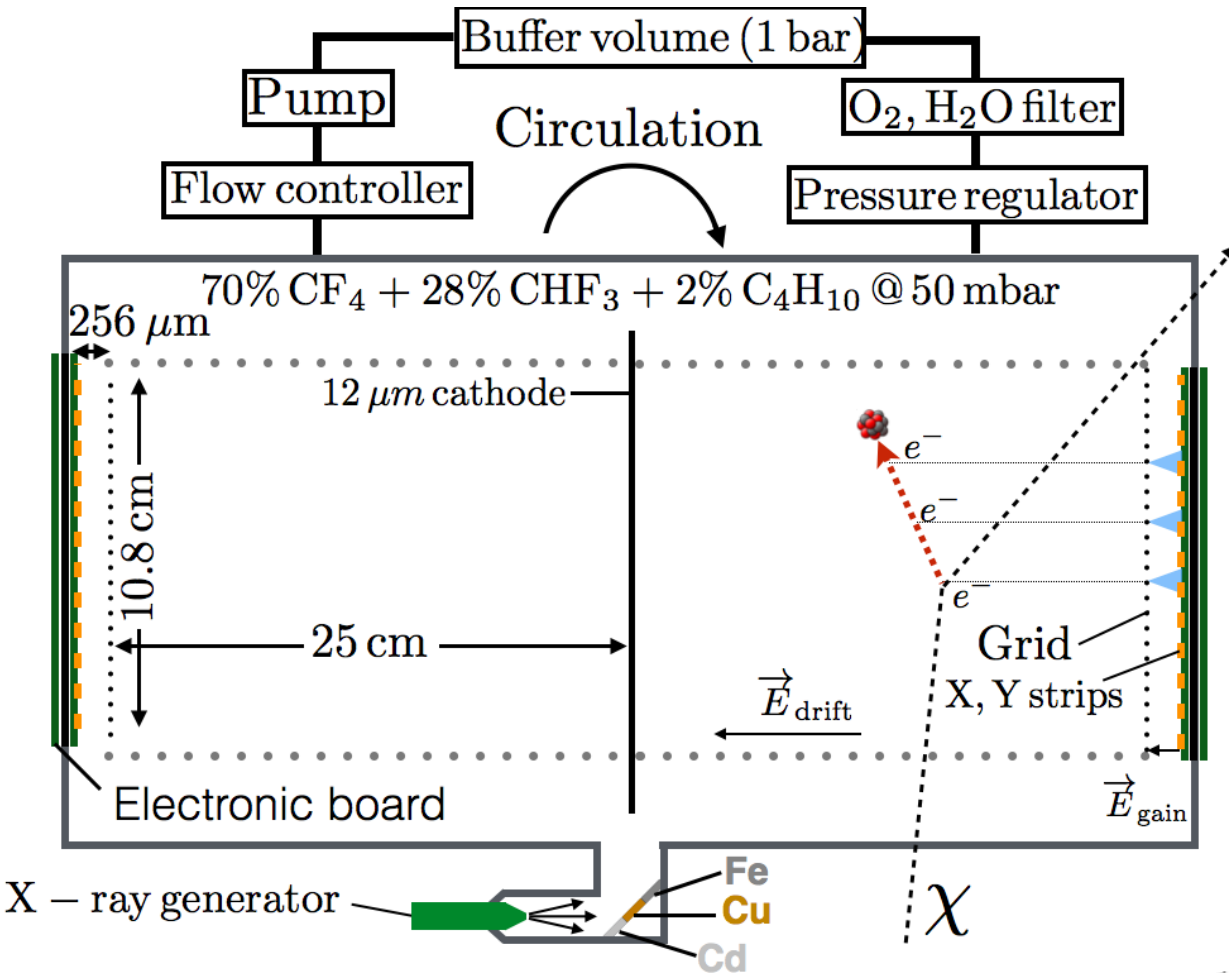
Scheme of a MIMAC μ TPC

Evolution of the collected charges on the anode

Measurement of the ionization energy:

Charge integrator connected to the mesh coupled to a FADC sampled at 50 MHz

MIMAC-bi-chamber module prototype



MIMAC Target: ^{19}F

- Light WIMP mass
- Axial coupling



MIMAC (bi-chamber module) at
Modane Underground Laboratory
(France)

since June 22nd 2012.

Upgraded in June 2013, and
in June 2014.

-working at 50 mbar
($\text{CF}_4 + 28\% \text{CHF}_3 + 2\% \text{C}_4\text{H}_{10}$)

-in a permanent circulating mode

-Remote controlled

and commanded

-Calibration control twice per week

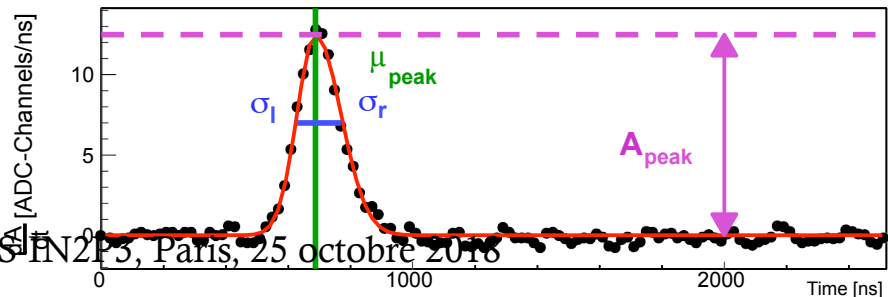
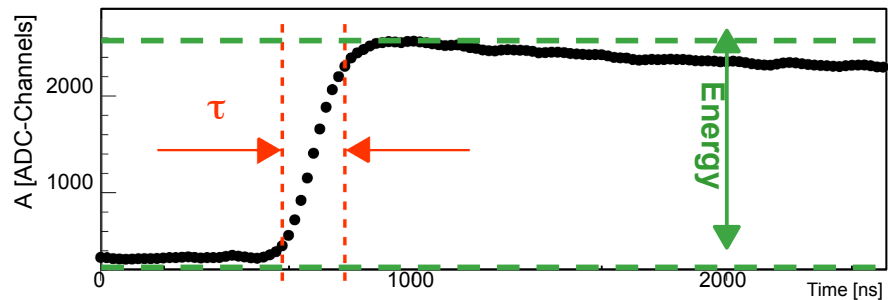
Many thanks to LSM staff

MIMAC readout

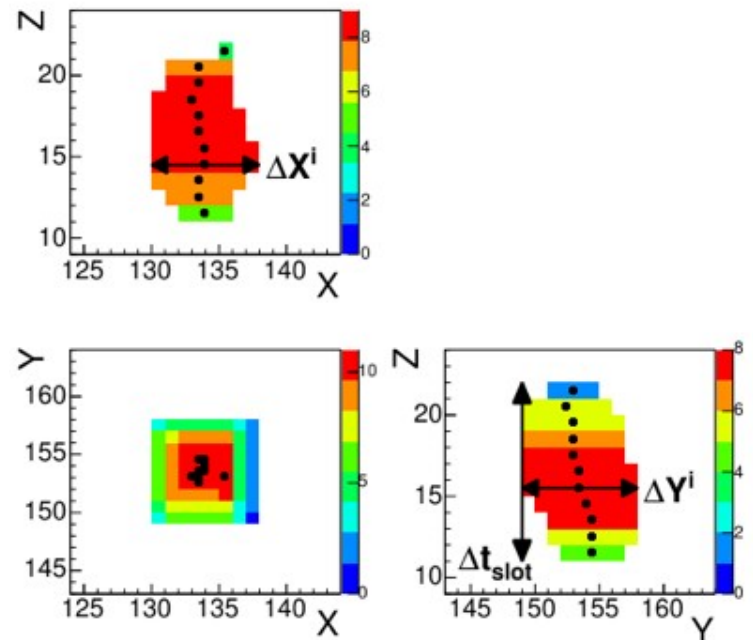


Dedicated fast electronics (self-triggered)
Based on the MIMAC chip (64 channels)

preamplifier signal + FADC: Energy



3D - track



Detector calibration (not at the maximum gain!)

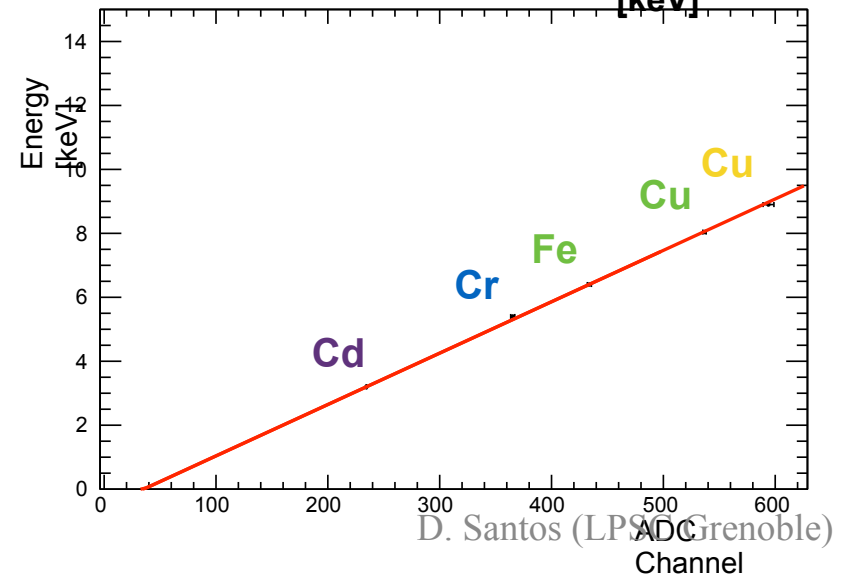
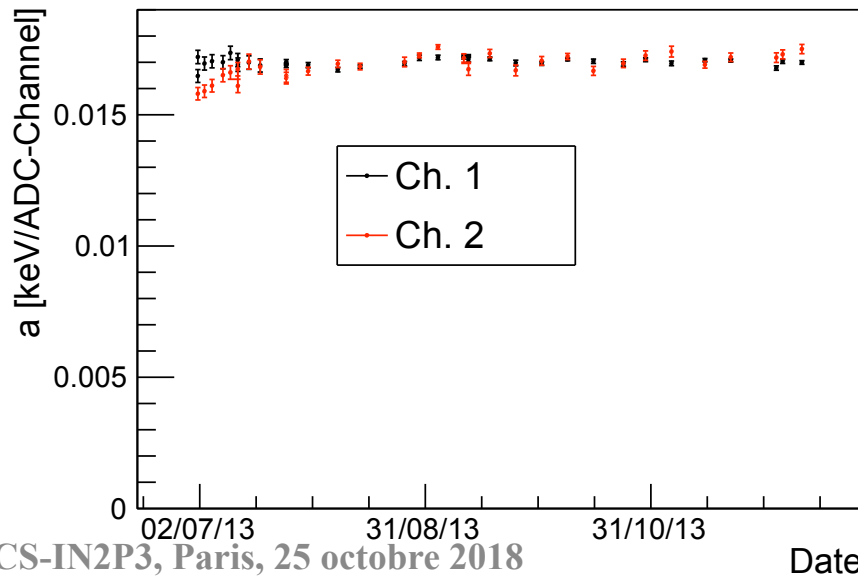
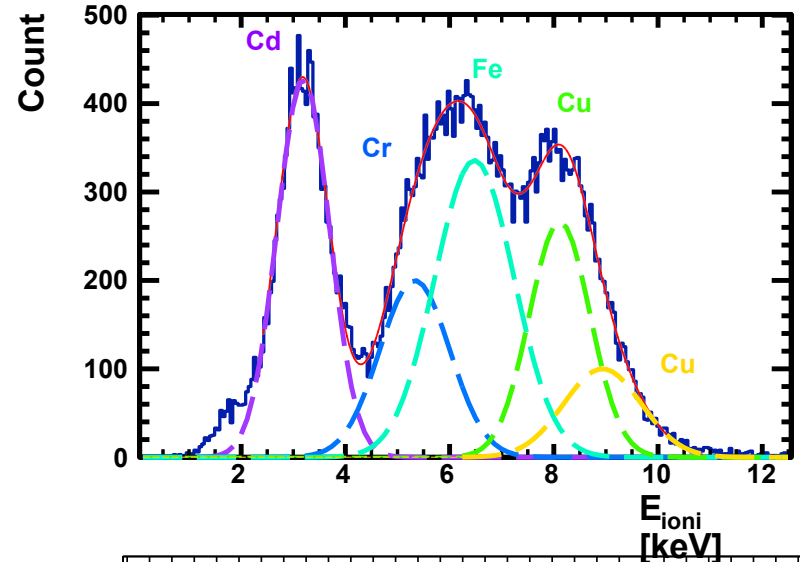
Calibration: (once a week)

X-ray generator producing fluorescence photons from Cd, Fe, Cu foils.

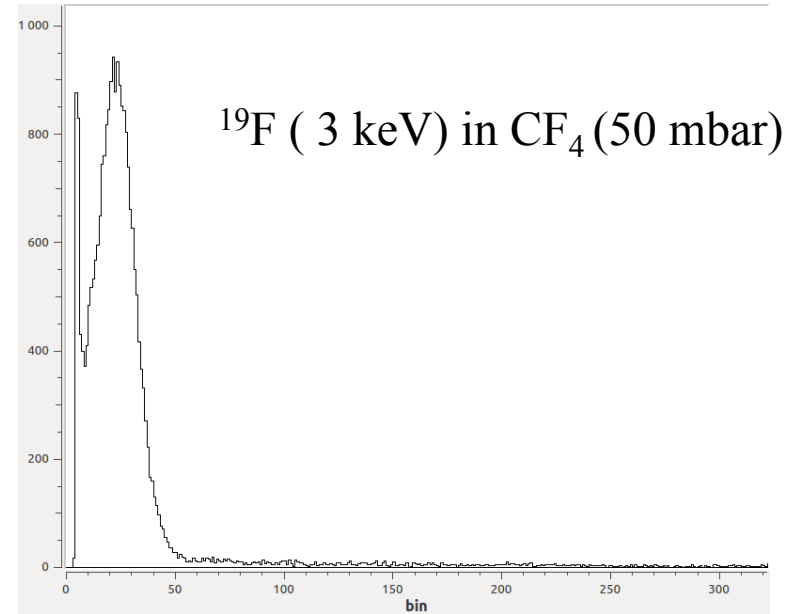
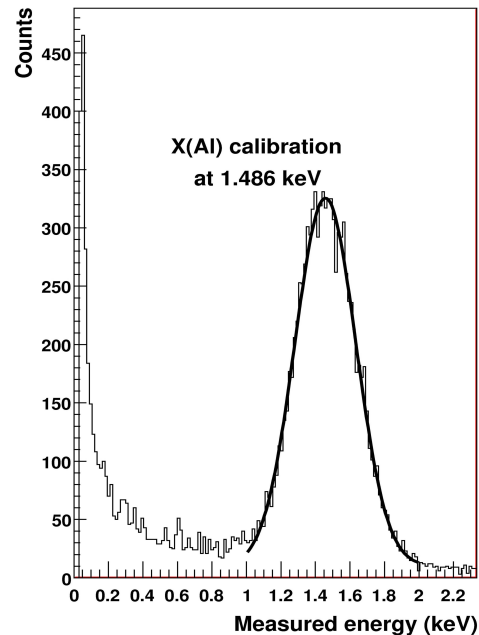
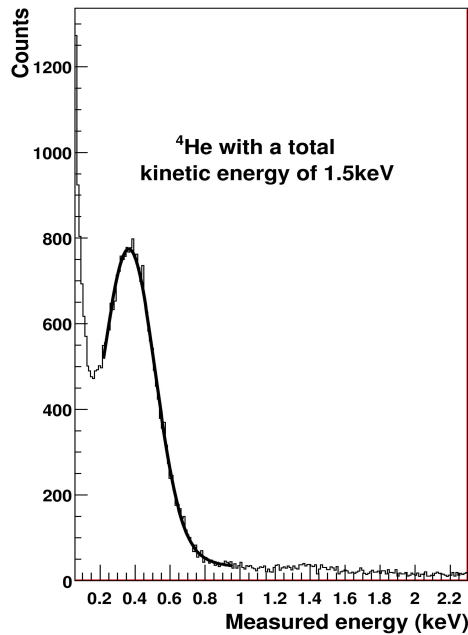
Threshold ~ 1 keV

Circulation system:

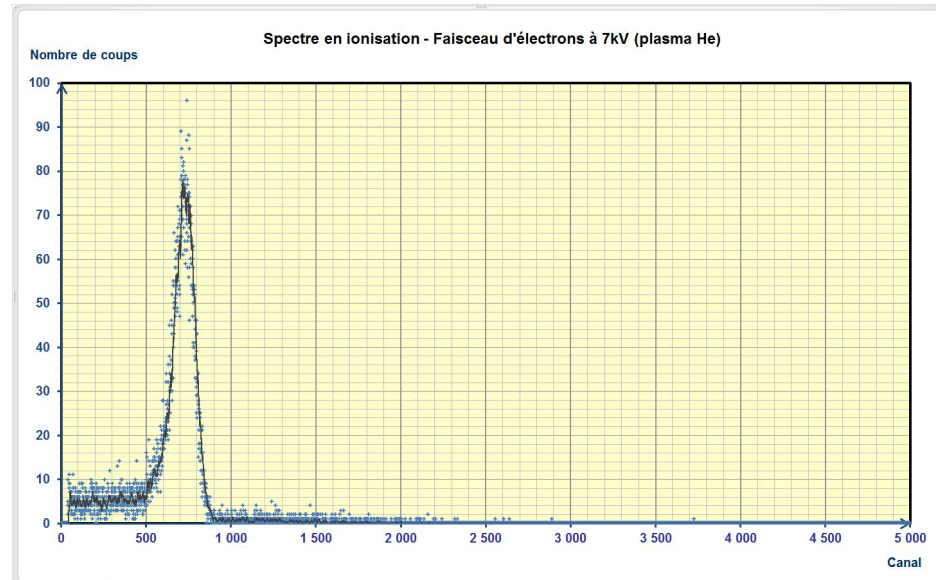
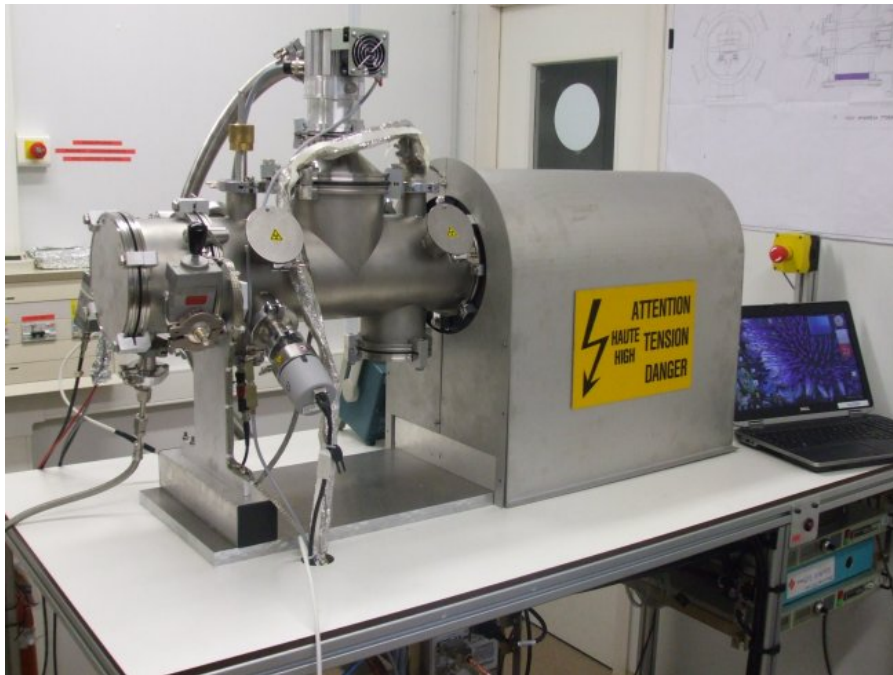
Excellent Gain stability in time



Ionization Quenching Factor Measurements at LPSC-Grenoble

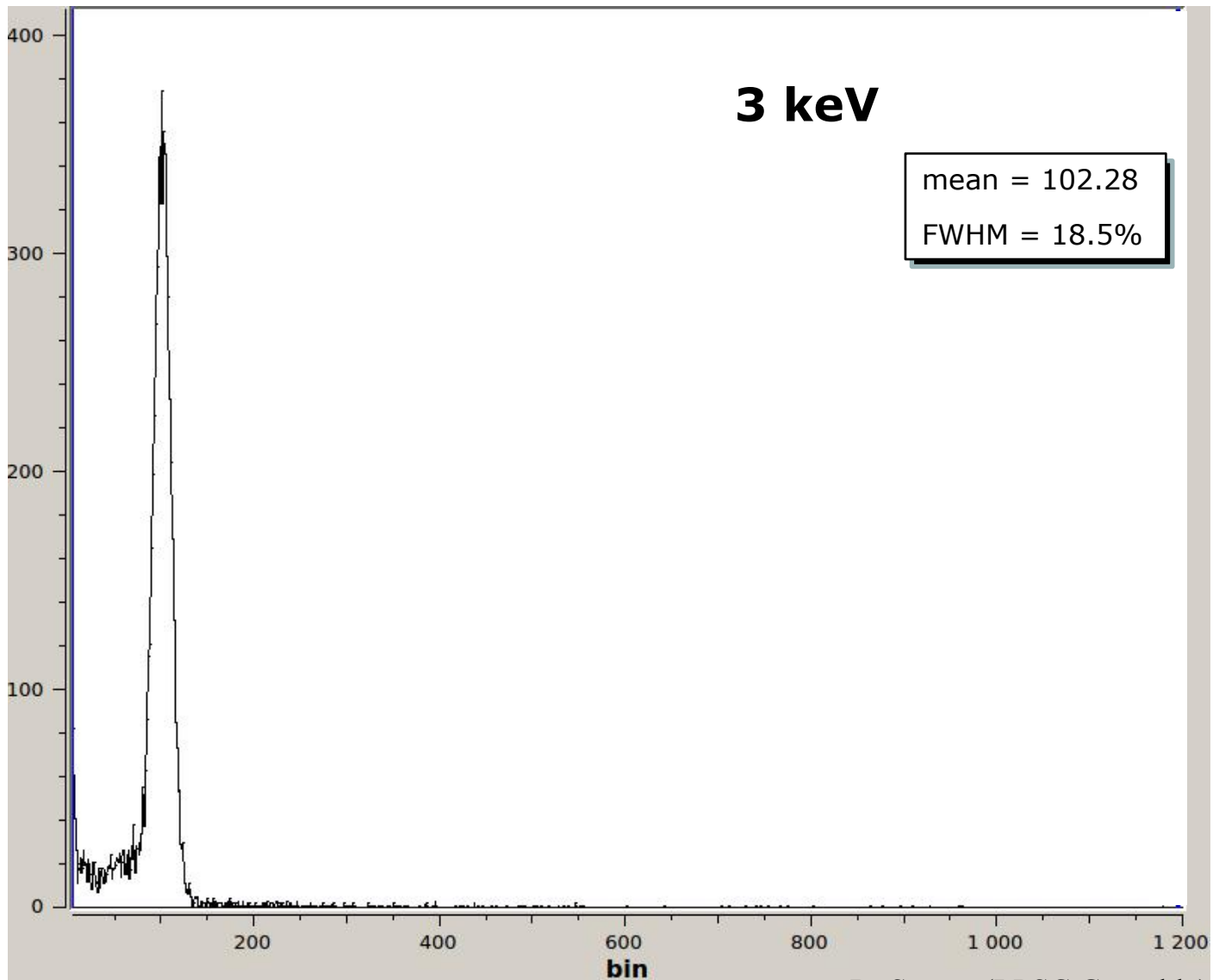


Portable Quenching Facility (COMIMAC) (Electrons and Nuclei of known energies)

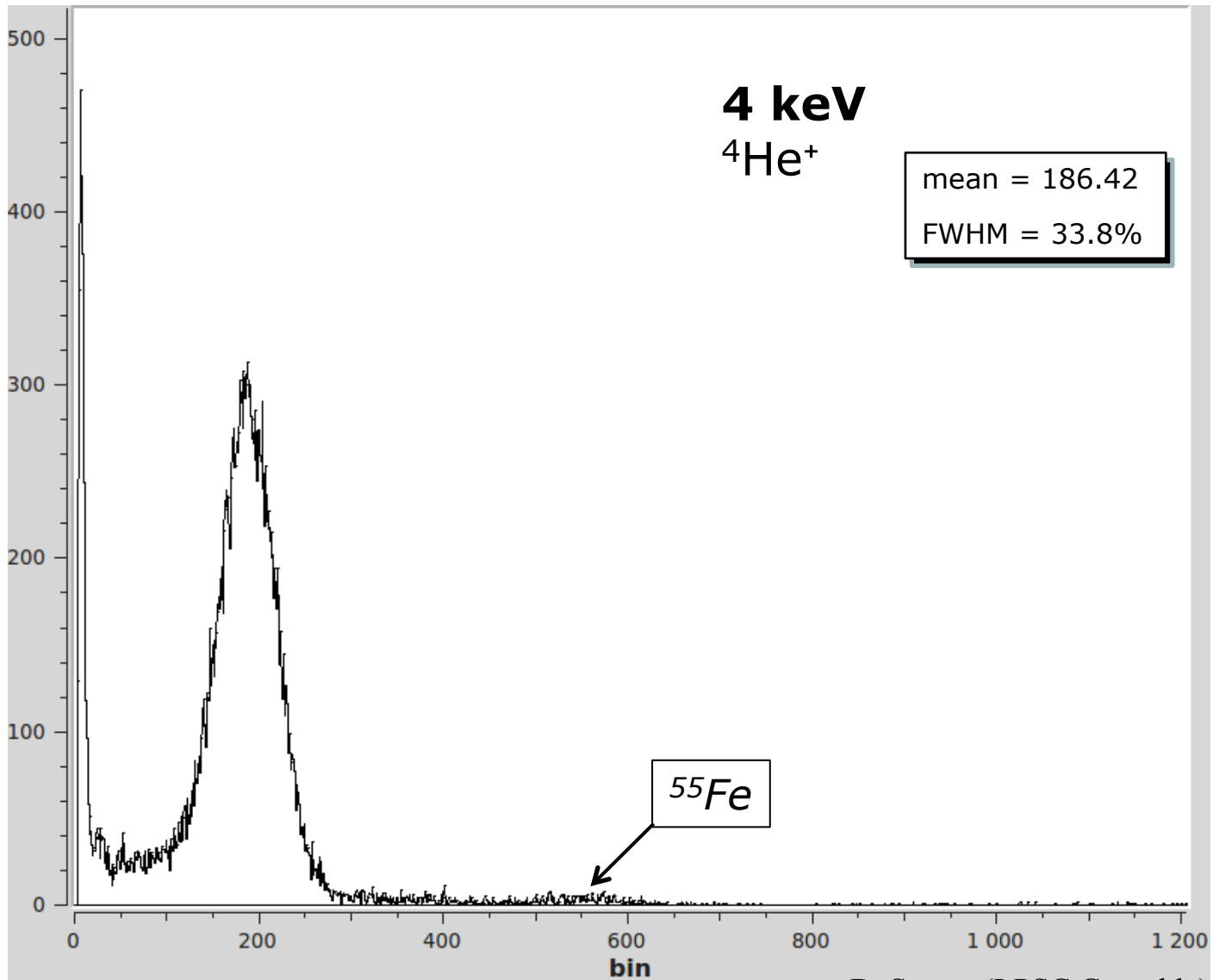


Electrons of 7 keV

**In a gas detector the IQF depends strongly on the quality of the gas.
The IQF needs to be measured periodically (in-situ) in a long term run experiment.**

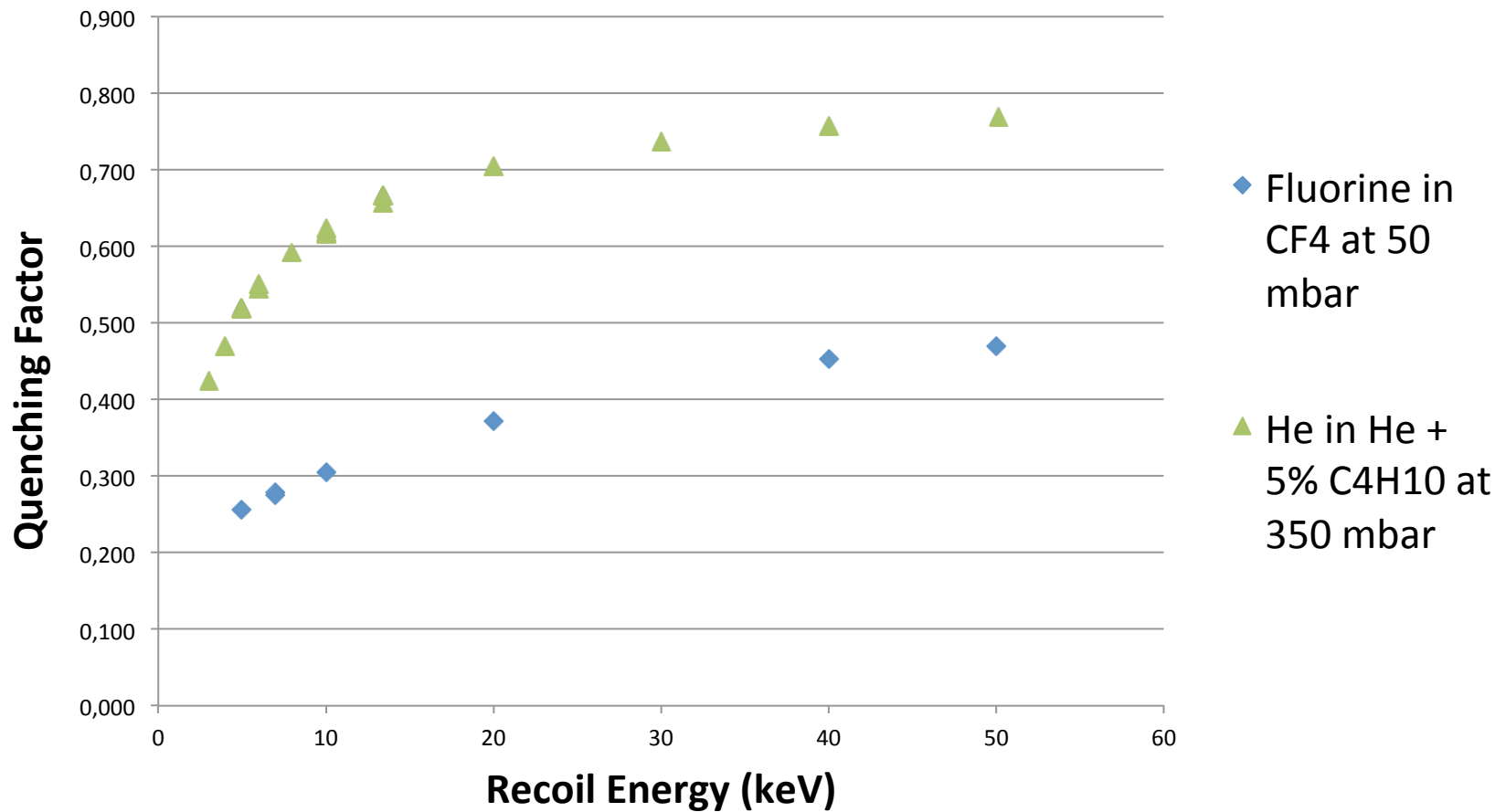


D. Santos (LPSC Grenoble)



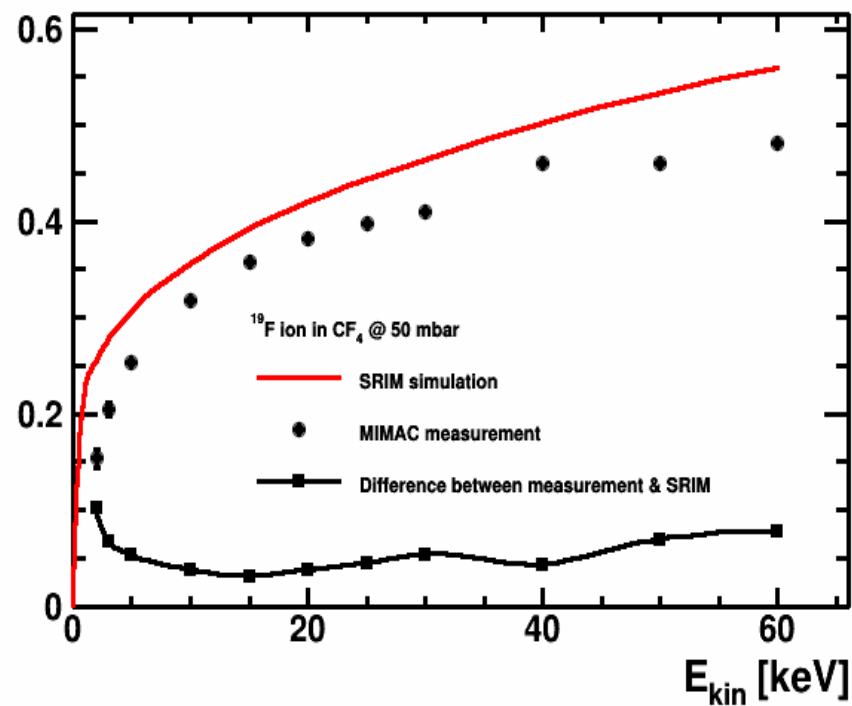
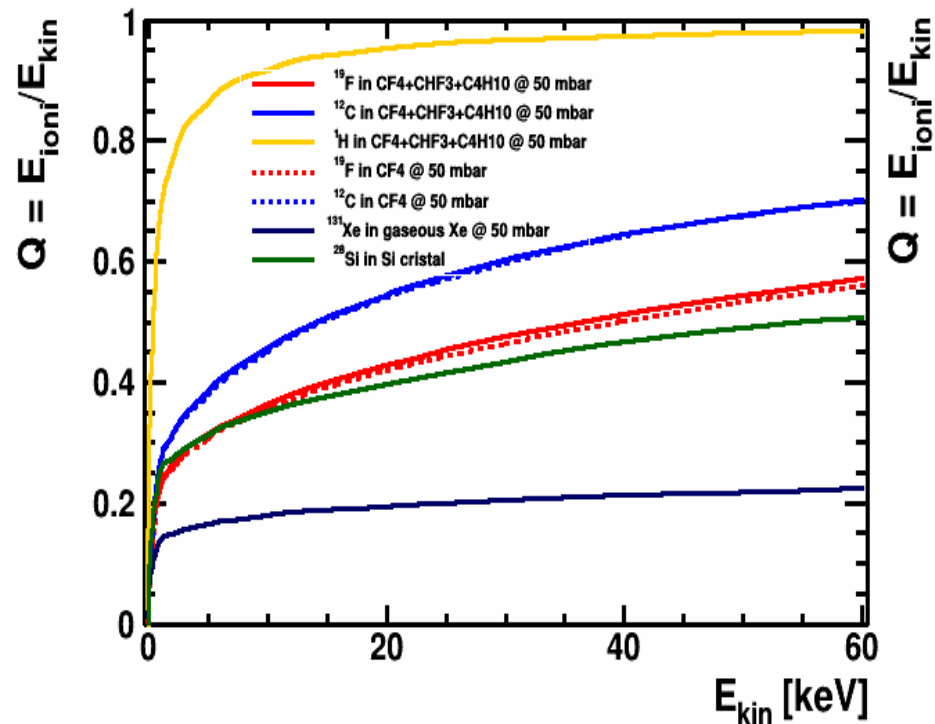
D. Santos (LPSC Grenoble)

Ionization Quenching Factor for Fluorine in pure CF4 at 50 mbar

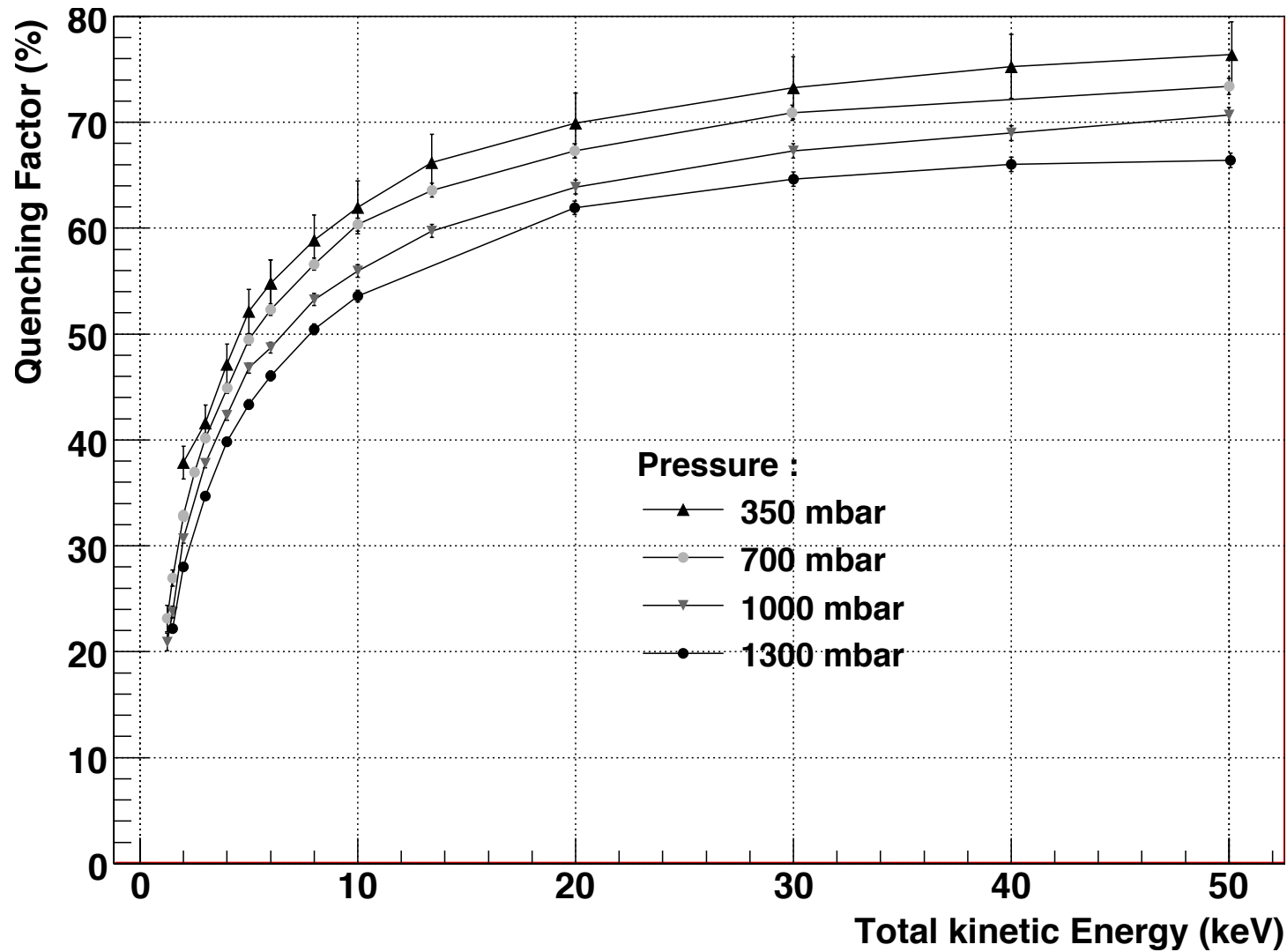


Ionization Quenching Factors

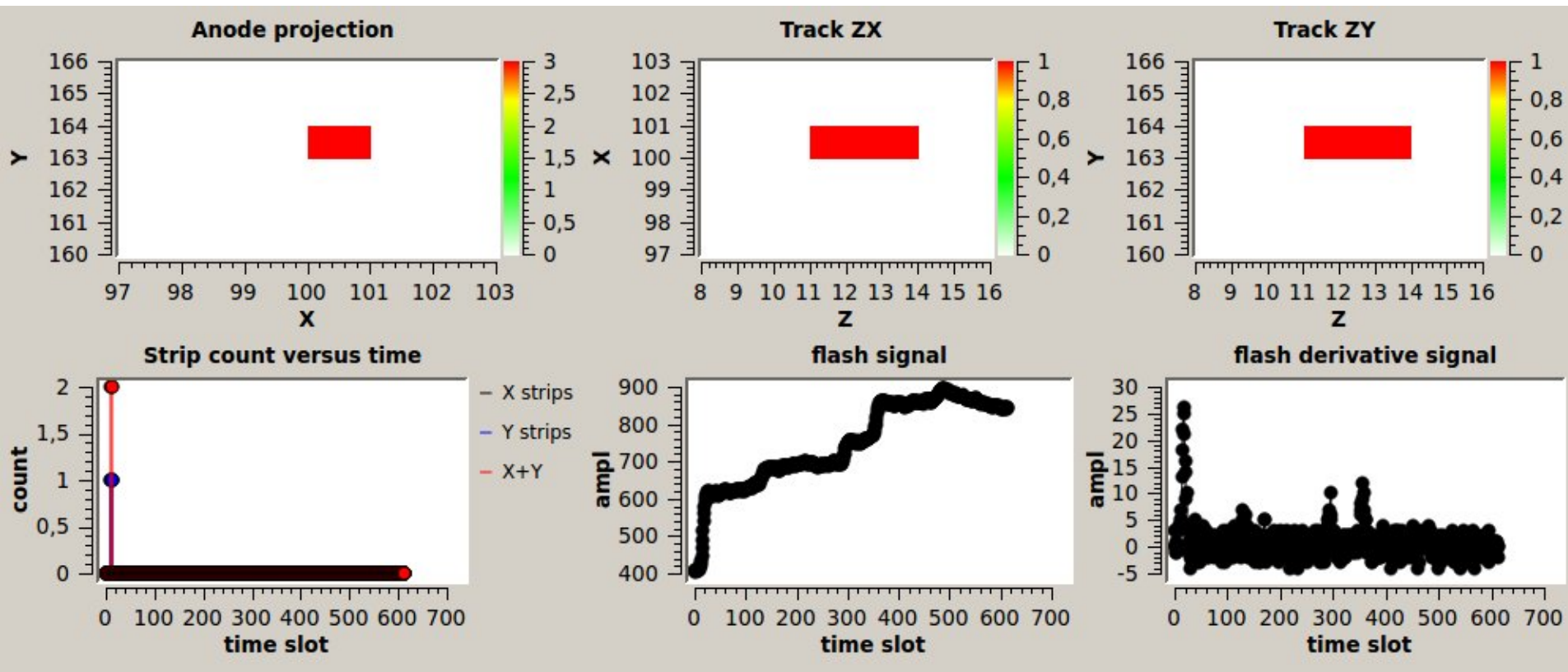
Simulations and Measurements (LPSC)



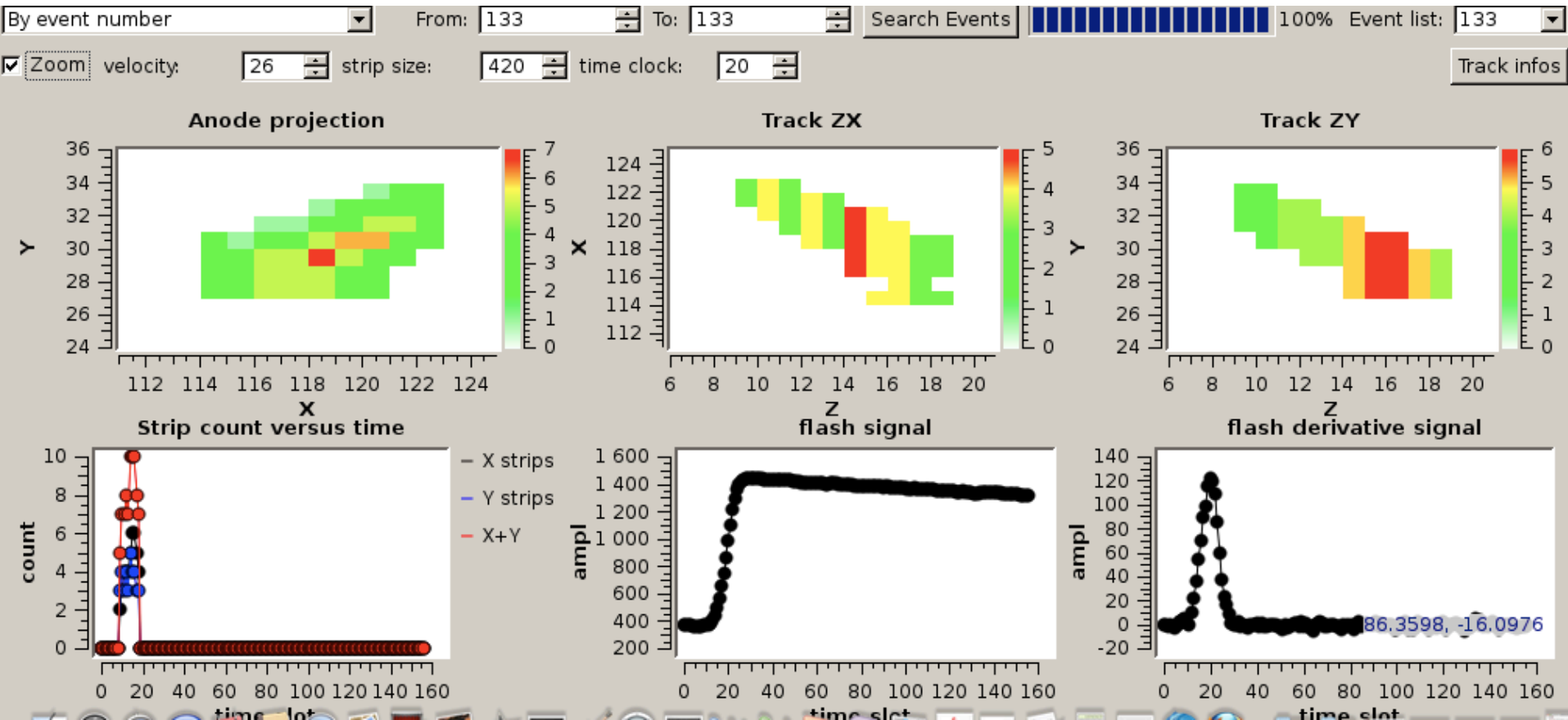
IQF in ^4He + 5% isobutane for different pressures!!



An Electron event (18 keV)

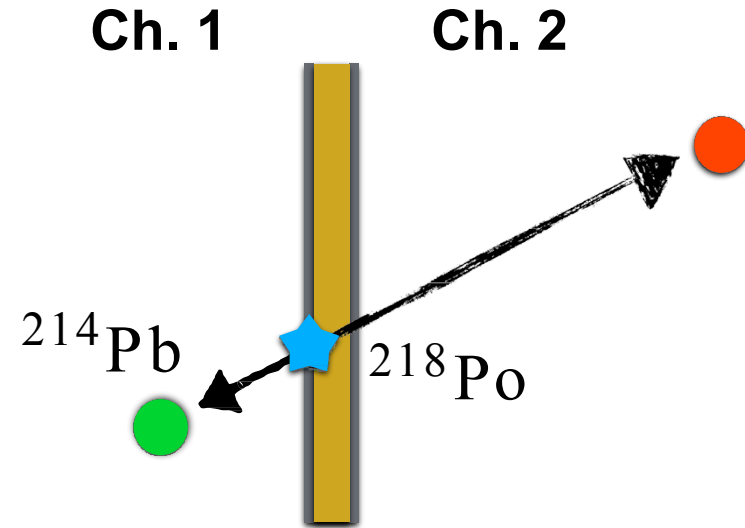
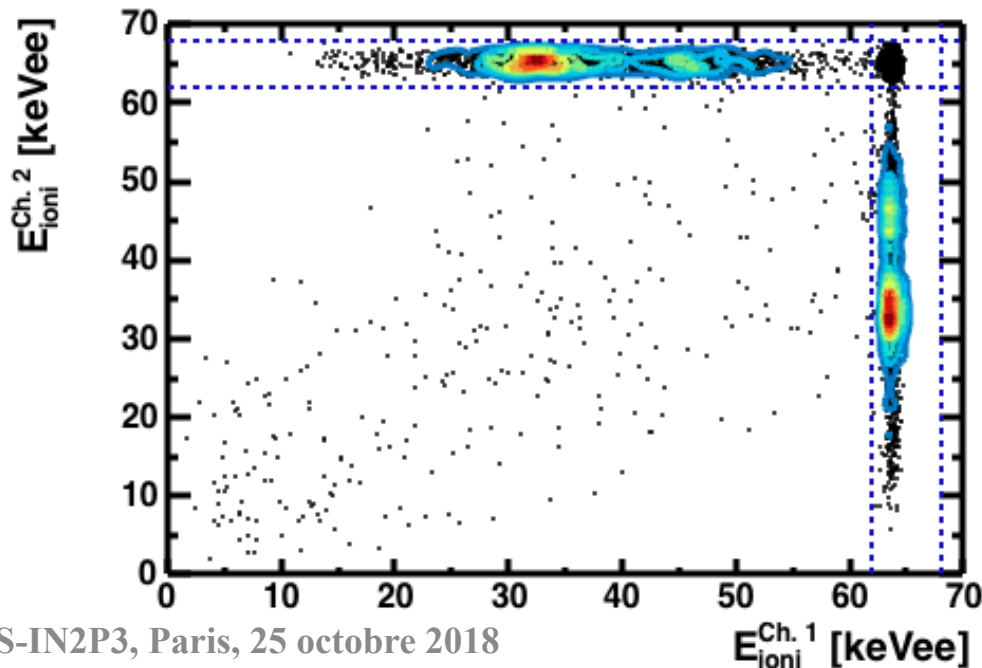


A “recoil event” (~ 34 keVee)



RPR: « In coincidence » events

Chamber coincidences:



3D tracks from nuclear recoil
of radon progeny detection

First detection of 3D tracks of Rn progeny

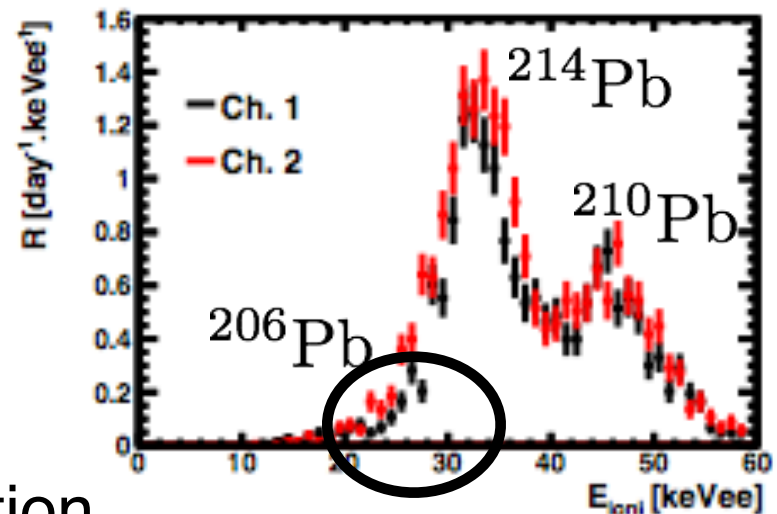
Electron/recoil discrimination

Mesure:
$$\begin{cases} E_{ioni}(^{214}\text{Pb}) = 32.90 \pm 0.16 \text{ keVee} \\ E_{ioni}(^{210}\text{Pb}) = 45.60 \pm 0.29 \text{ keVee} \end{cases}$$

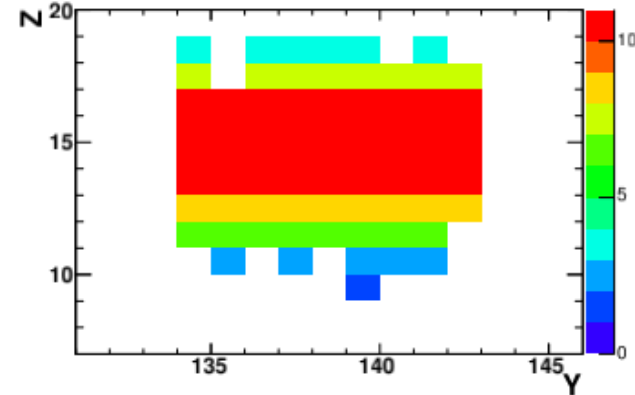
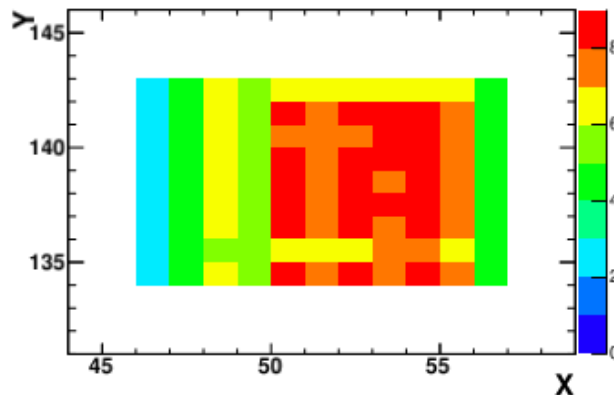
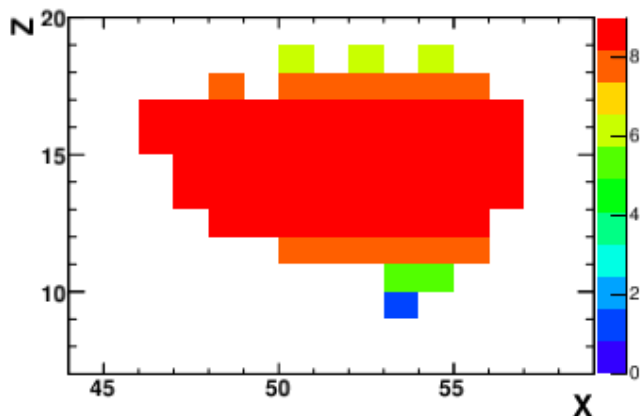
First measurement of 3D nuclear-recoil tracks coming from radon progeny

→ MIMAC detection strategy validation

Nuclear recoil spectra



$$R_{206\text{Pb}} \sim 0.25 \text{ day}^{-1} \cdot \text{keVee}^{-1}$$

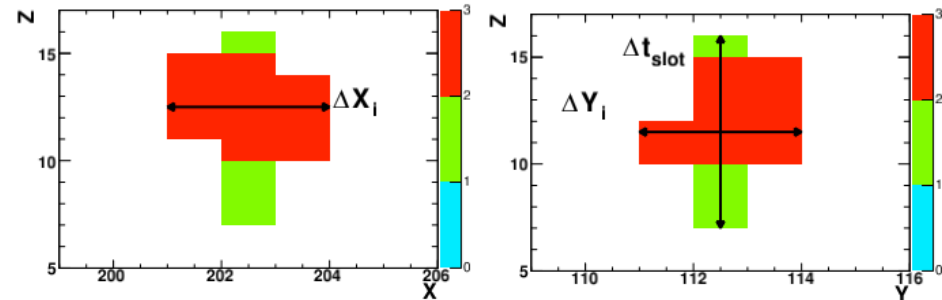


RPR events occur at different positions in the detector...

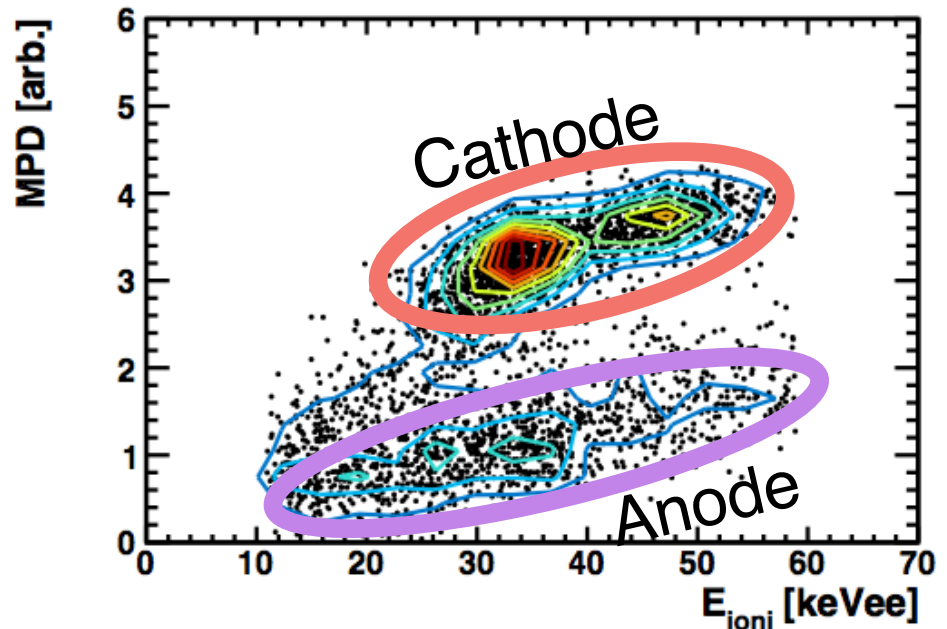
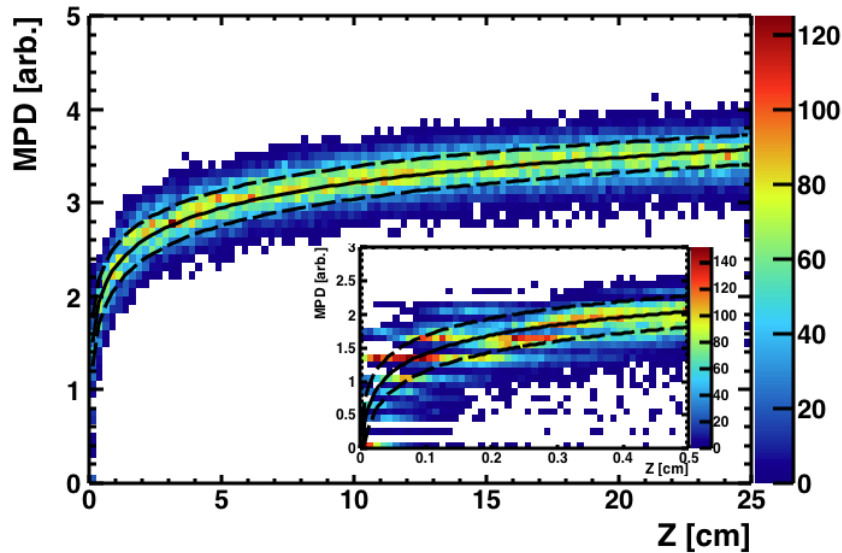
$z_0 \longleftrightarrow$ Diffusion

$$\begin{cases} D_T = 237.9 \mu\text{m}/\sqrt{\text{cm}} \\ D_L = 271.5 \mu\text{m}/\sqrt{\text{cm}} \end{cases}$$

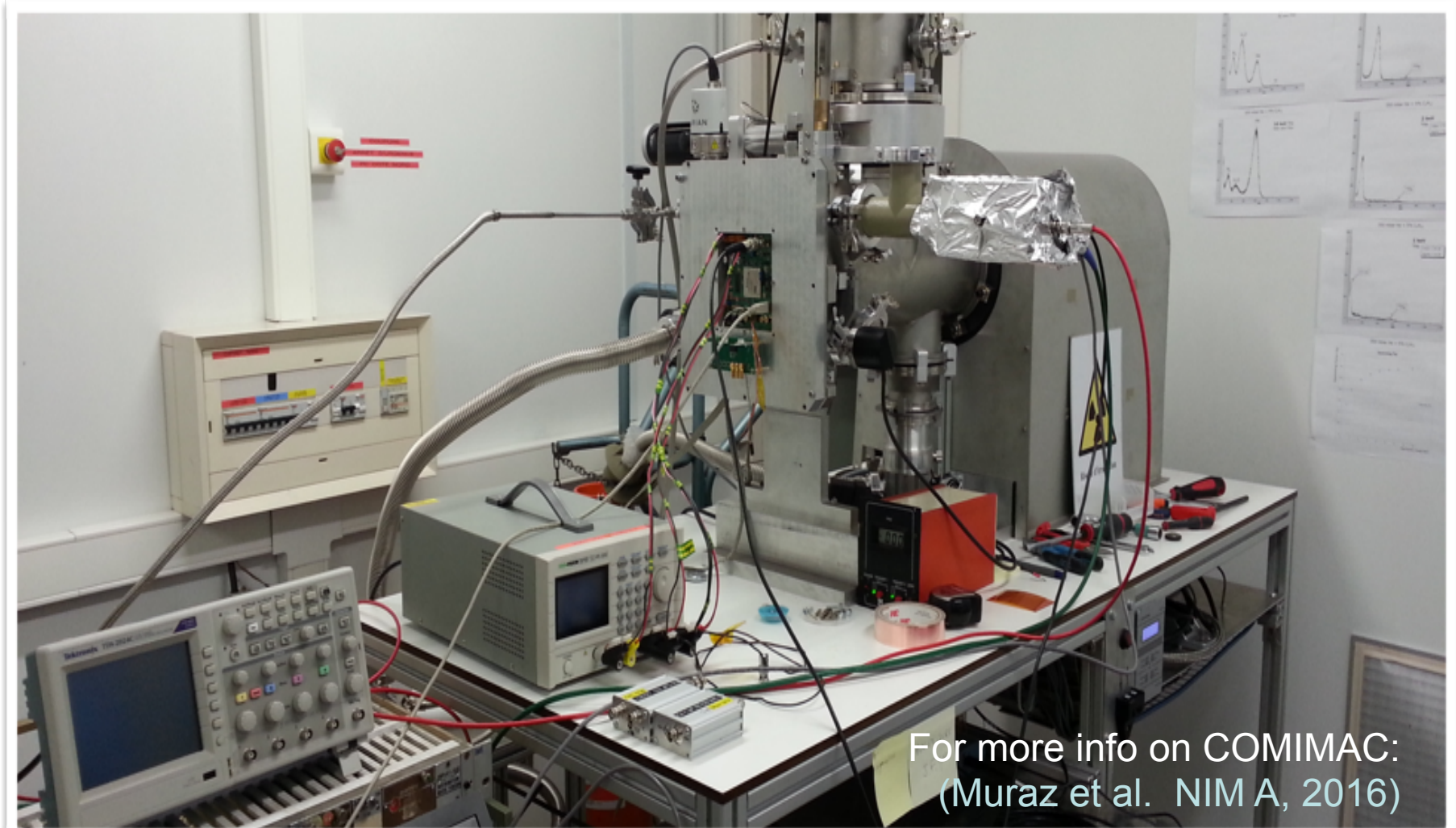
« Anode » event



Mean Projected Diffusion: $\bar{D} = \ln(\overline{\Delta X} \times \overline{\Delta Y})$



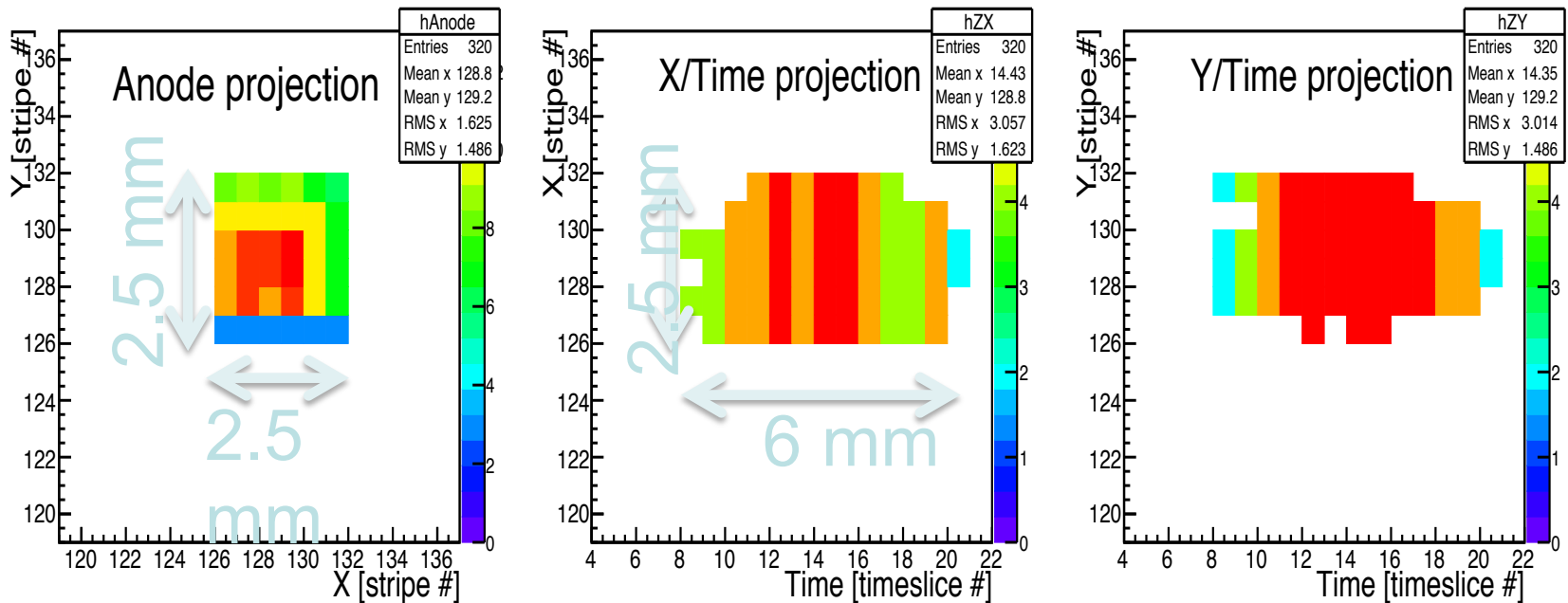
First controlled Fluorine tracks, using COMIMAC



For more info on COMIMAC:
(Muraz et al. NIM A, 2016)

COMIMAC: first measurements on controlled tracks of Fluorine

25 keV (kinetic) Fluorine \rightarrow \sim 9 keVee

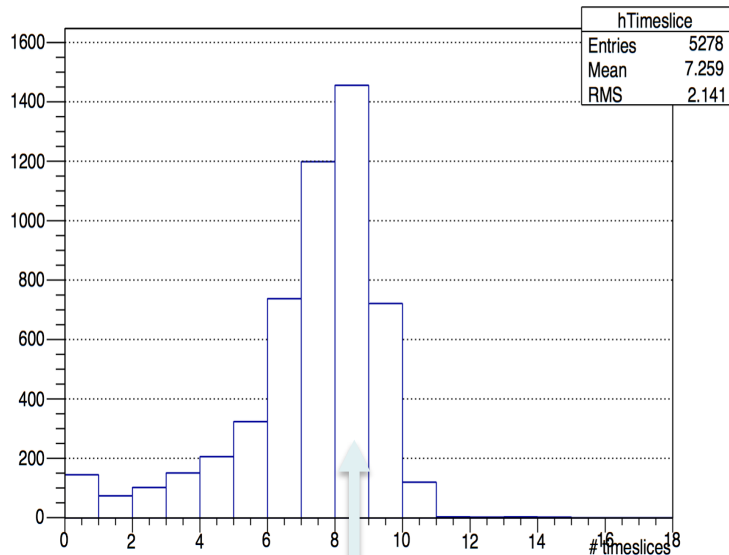


D. Santos (LPSC Grenoble)

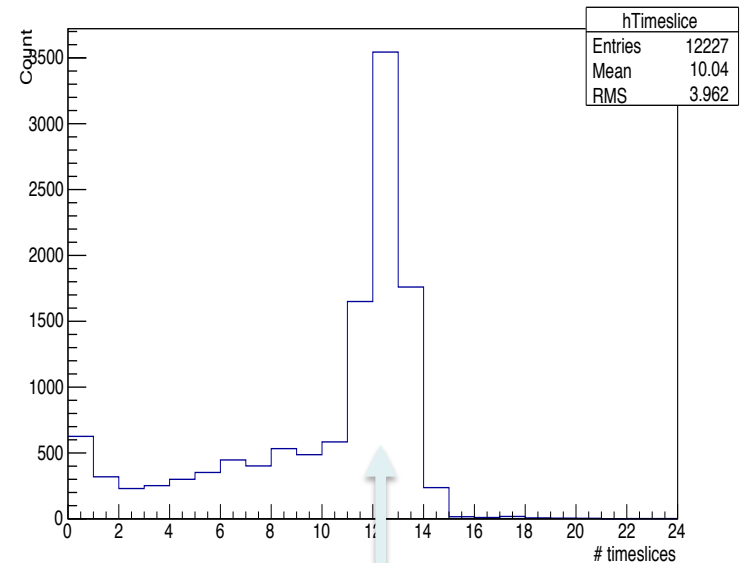
COMIMAC: first controlled tracks of ^{19}F

8 keV kinetic \rightarrow 2 keVee

25 keV kinetic \rightarrow 9 keVee



8 timeslices
* 20 ns/timeslices
* 23.5 $\mu\text{m}/\text{ns}$
= 3.8 mm



12 timeslices
* 20 ns/timeslice
* 23.5 $\mu\text{m}/\text{ns}$
= 5.8 mm

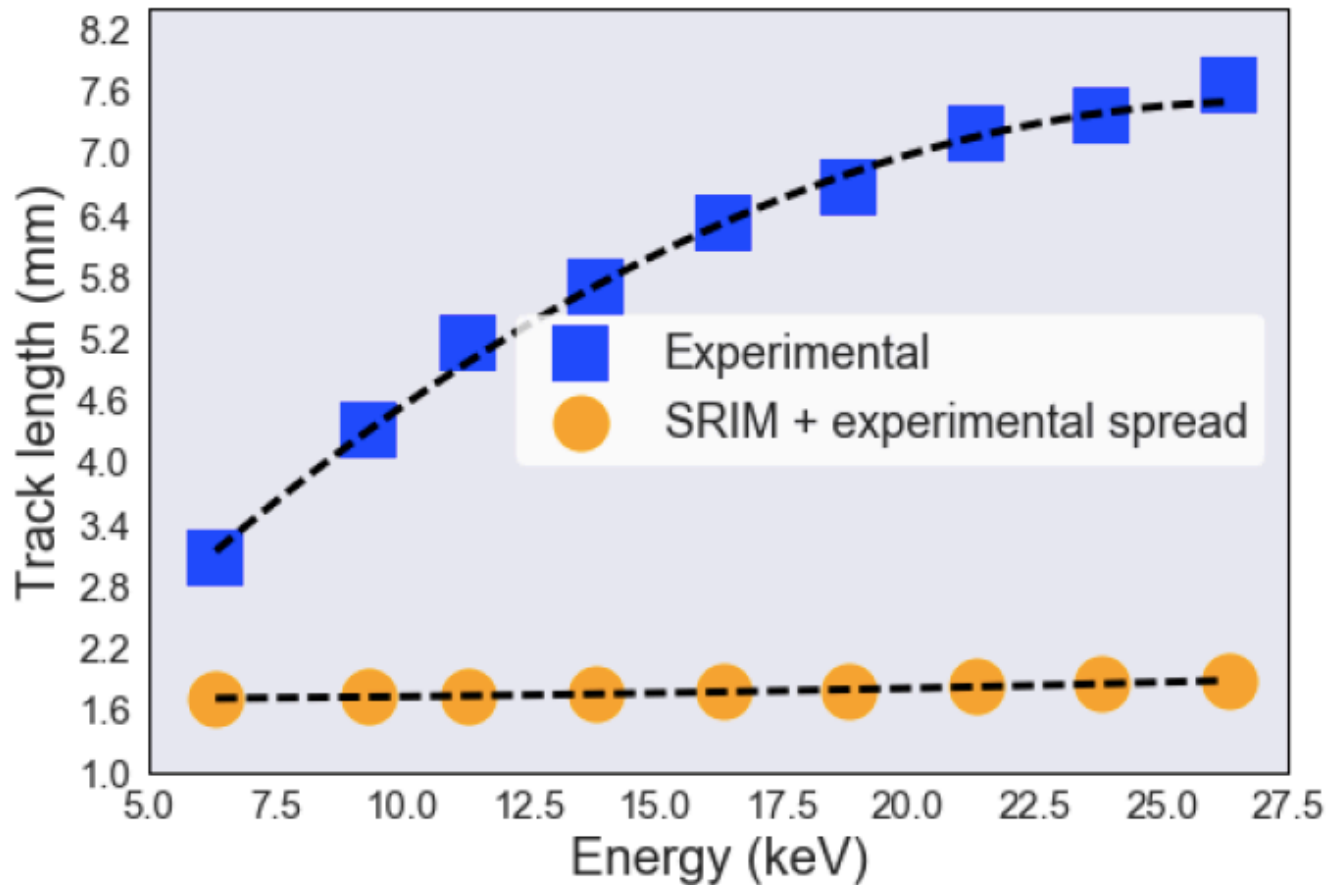
I. Moric, Y. Tao et al. (in preparation)

CS-IN2P3, Paris, 25 octobre 2018

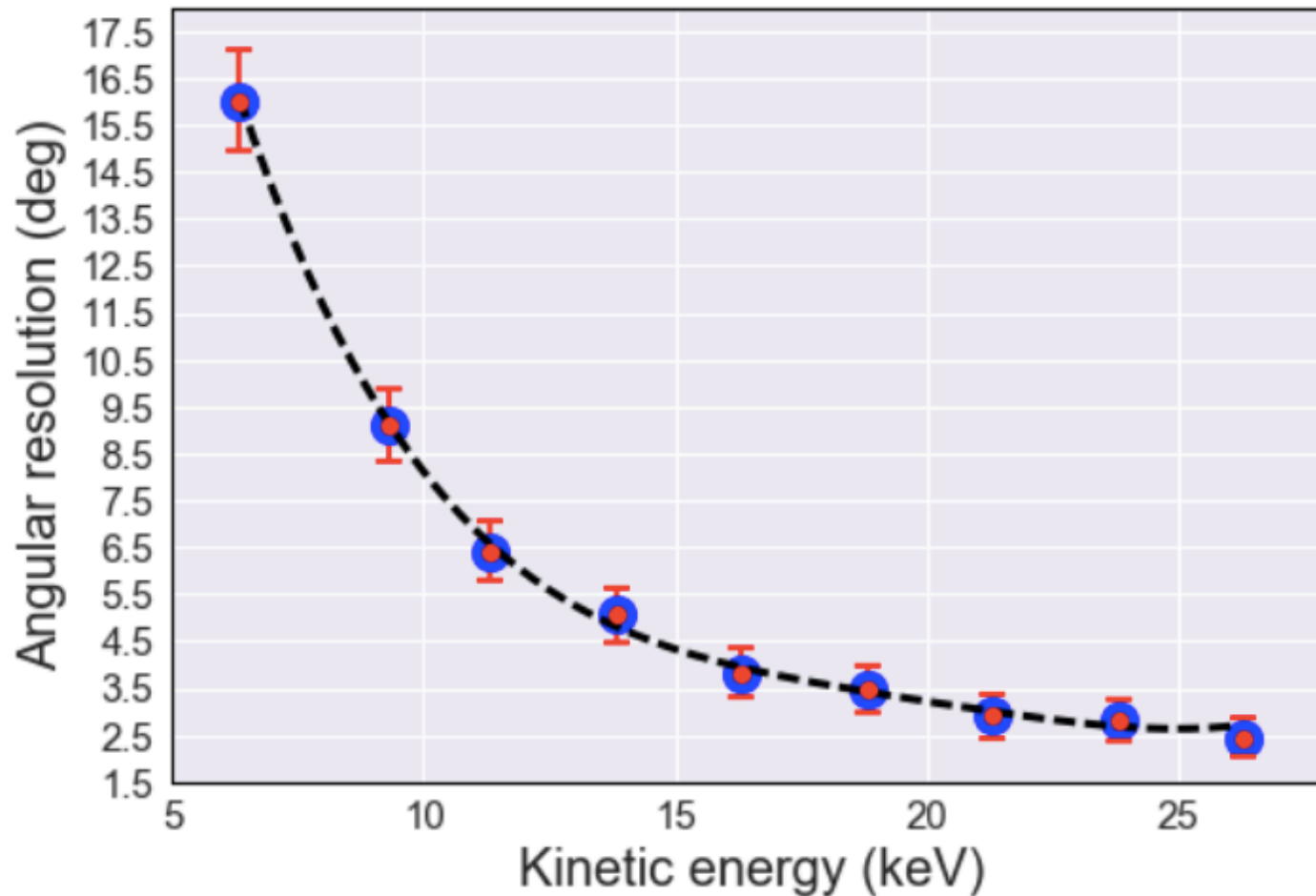
D. Santos (LPSC Grenoble)

3D-Track Length Measured and Simulated

(I. Moric, Y. Tao, N. Sauzet, DS. et al. 2018)

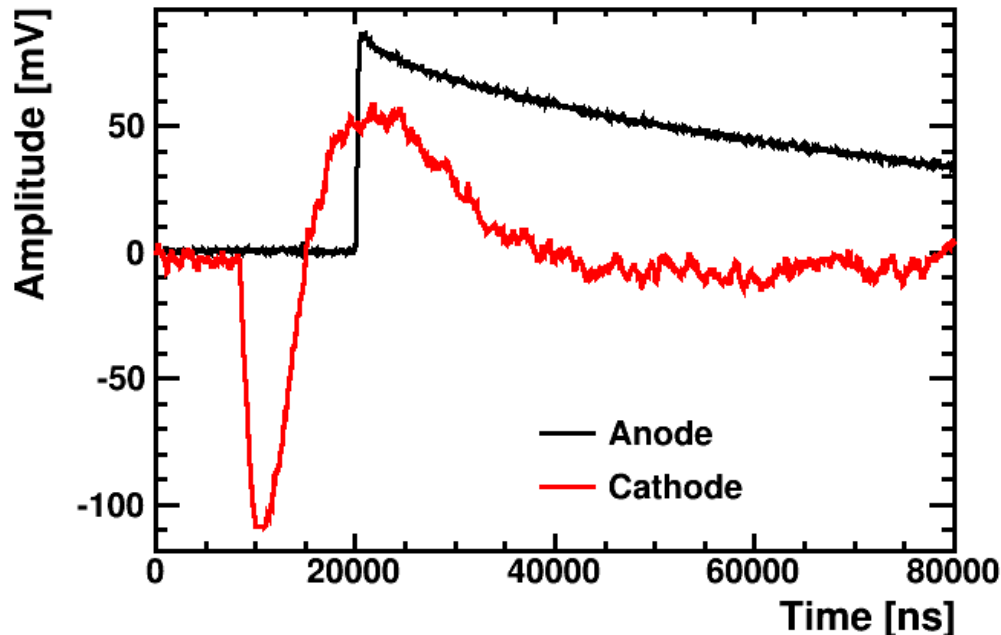


Angular resolution measured with COMIMAC
(^{19}F ions at known kinetic energies)
(I. Moric, Y. Tao et al. 2018)



Cathode Signal to place the 3D-track

- The cathode signal is produced by the primary electrons. It is produced before the anode signal produced by the avalanche.



Measurement in a MIMAC chamber of an alpha passing through the active volume parallel to the cathode at 10 cm distance.

MIMAC-Cathode Signal measurements

C. Couturier, Q. Riffard, N. Sauzet, O. Guillaudin, F. Naraghi, and D. Santos.
Journal of Instrumentation, 12(11):P11020, 2017.

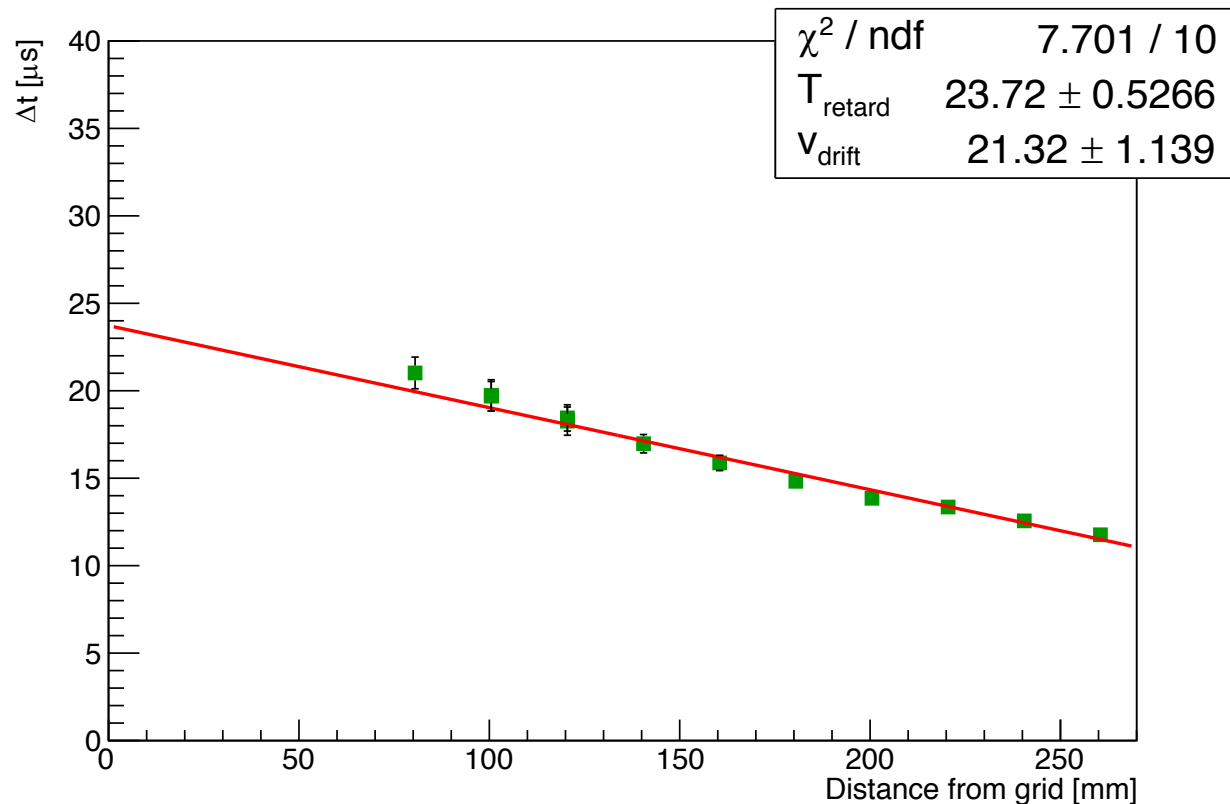
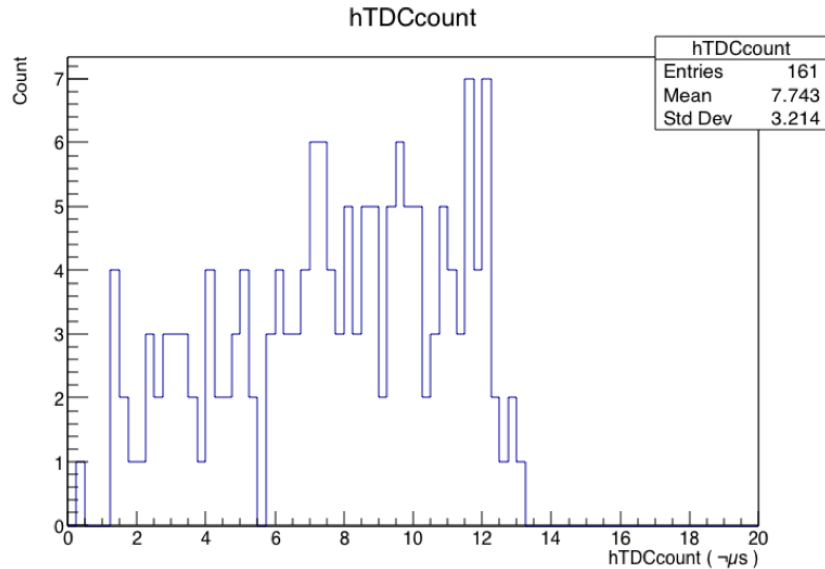


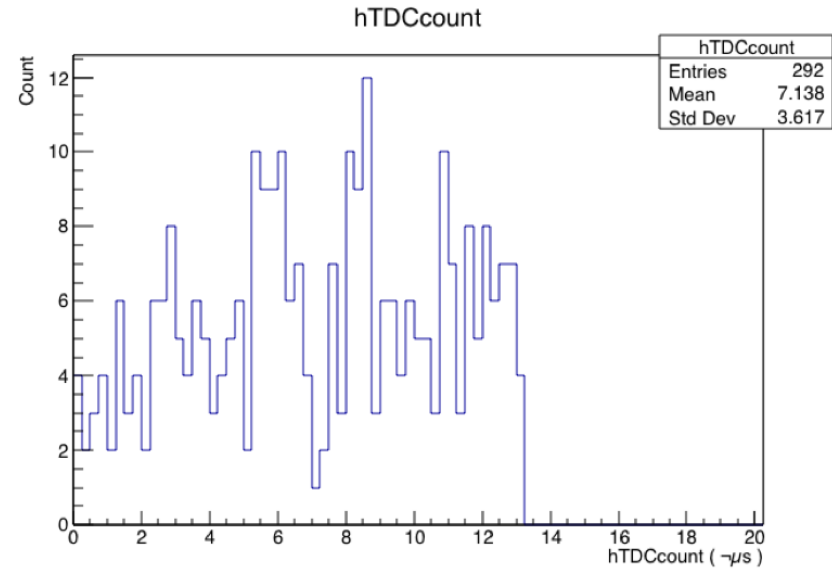
Figure 4. Measure of the time differences (TAC) between the grid signal and the delayed cathode signal in the “START Grid” configuration, as a function of the distance of the α source from the anode (green points) ; error bars correspond to the standard deviation of the mean. A linear fit of these points is superimposed in red and provides the values of the drift velocity and the additional delay.

First Cathode Signals from the MIMAC bichamber background (O. Guillaudin et al. October 2018)

Chamber 1

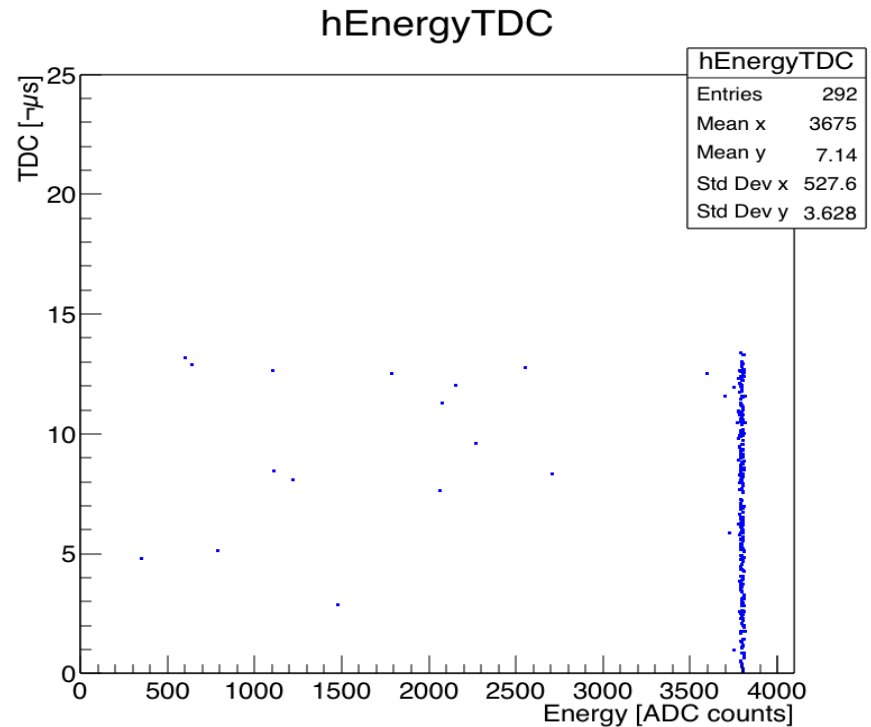
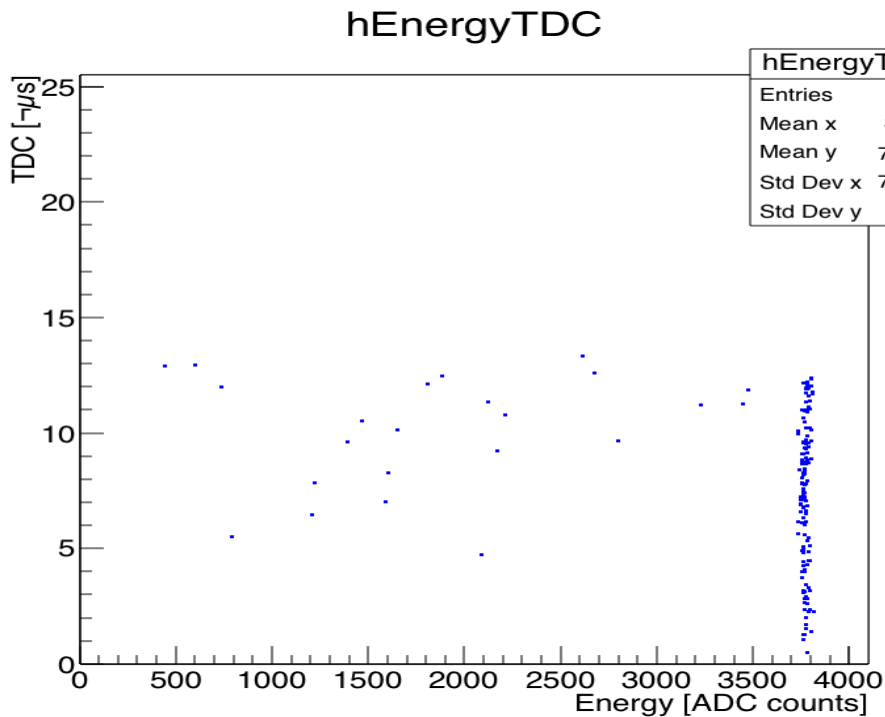


Chamber 2

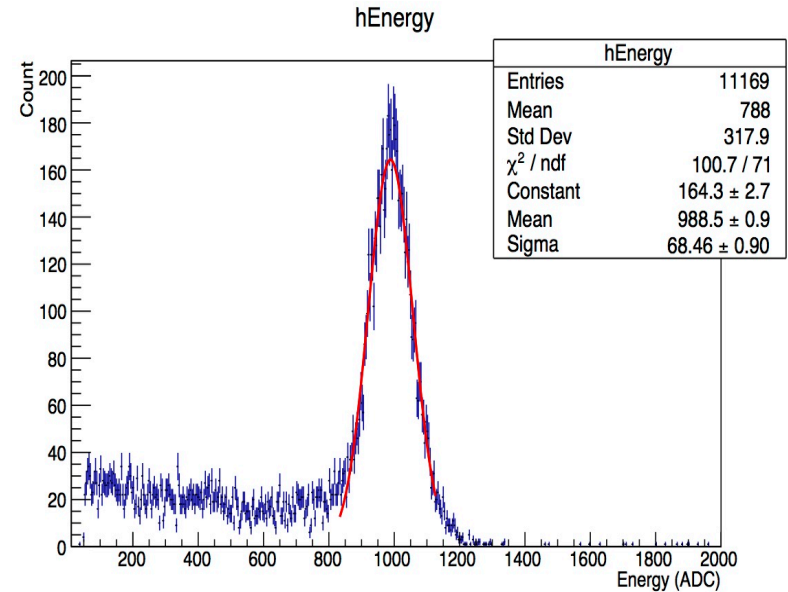
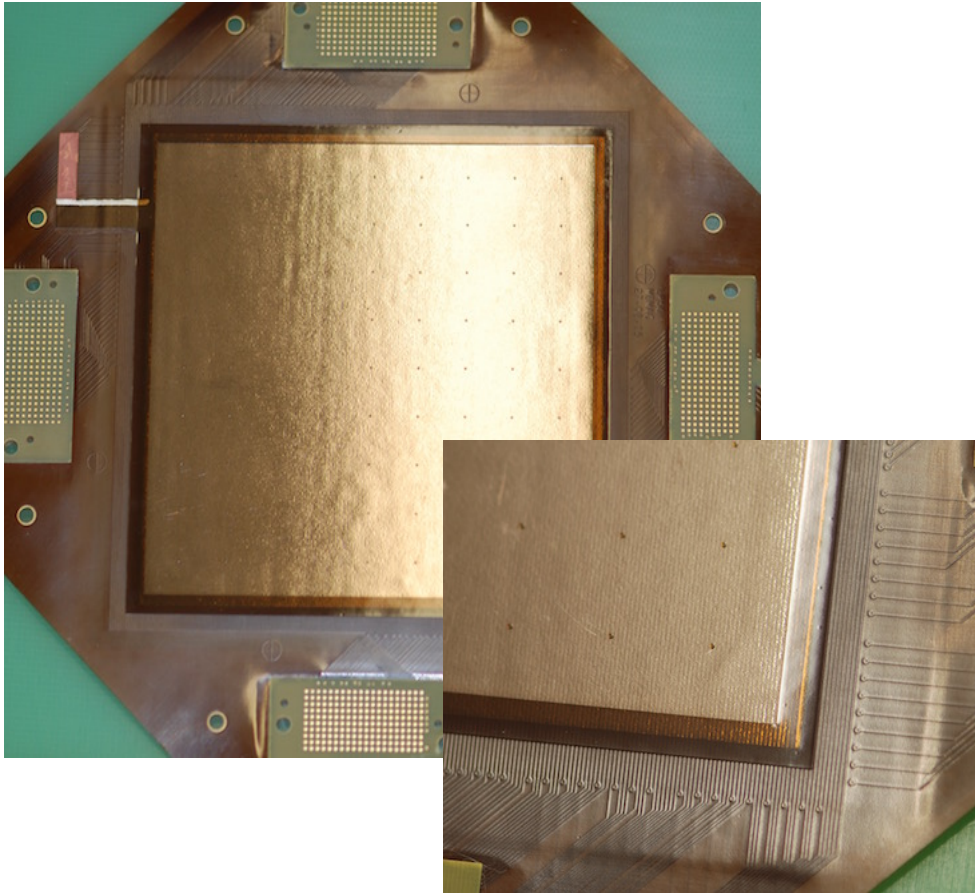


Measuring the time between the “event production” and the avalanche signal !!
Covering the 26 cm drift distance (13 us x 20 um/ns) !!

Ionization Energy distribution of the events recorded with the Cathode Signal



New MIMAC low background detector 10 cm x 10 cm

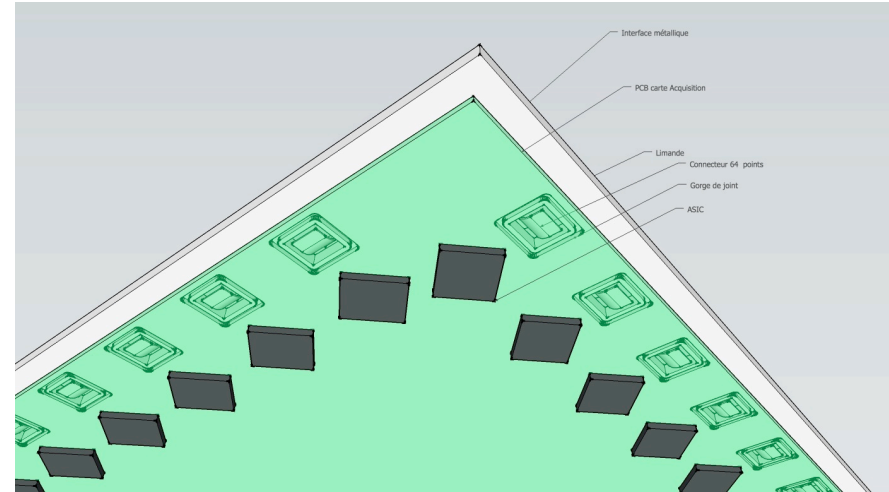
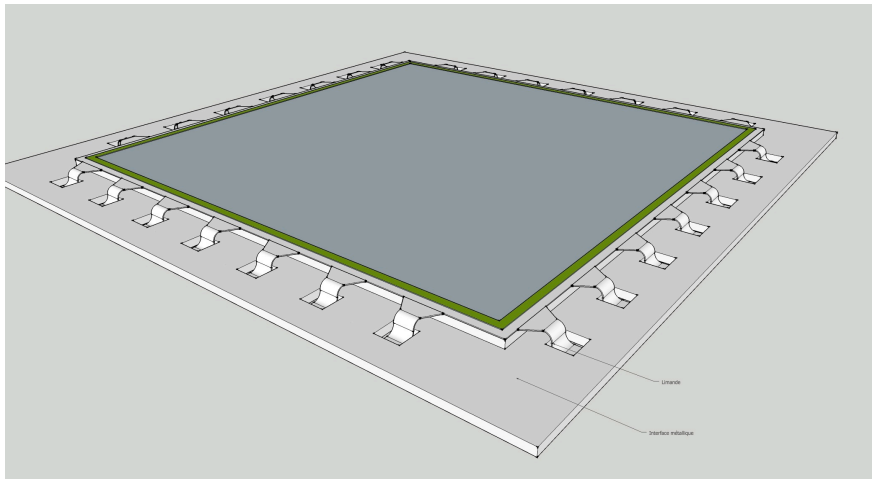


Gaz : MIMAC 50 mbar
HT grille : -560 V
Drift field : -150 V/cm

Kapton micromegas readout
Piralux Pilar

16,3 % FWHM (6 keV)
Gain ~25 000
Energy threshold <1 keV
D. Santos (LPSC Grenoble)

New 35 x 35 cm² low background detector design (1792 channels, 28 ASICs) (O. Guillaudin et al.)

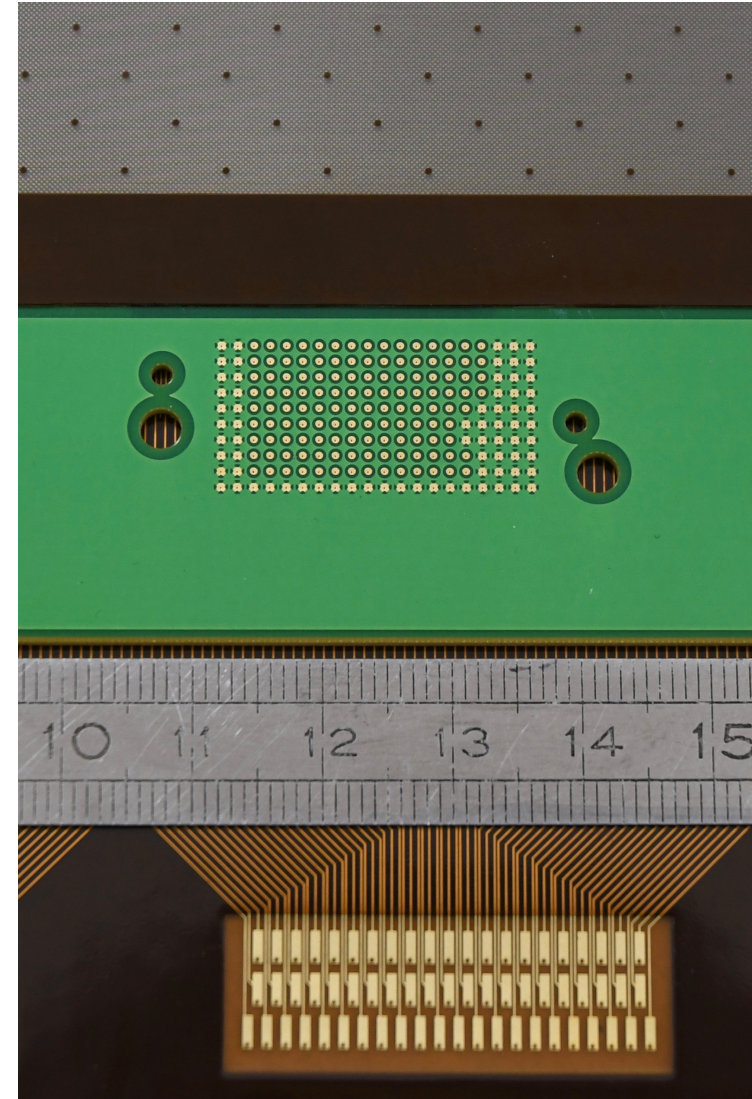
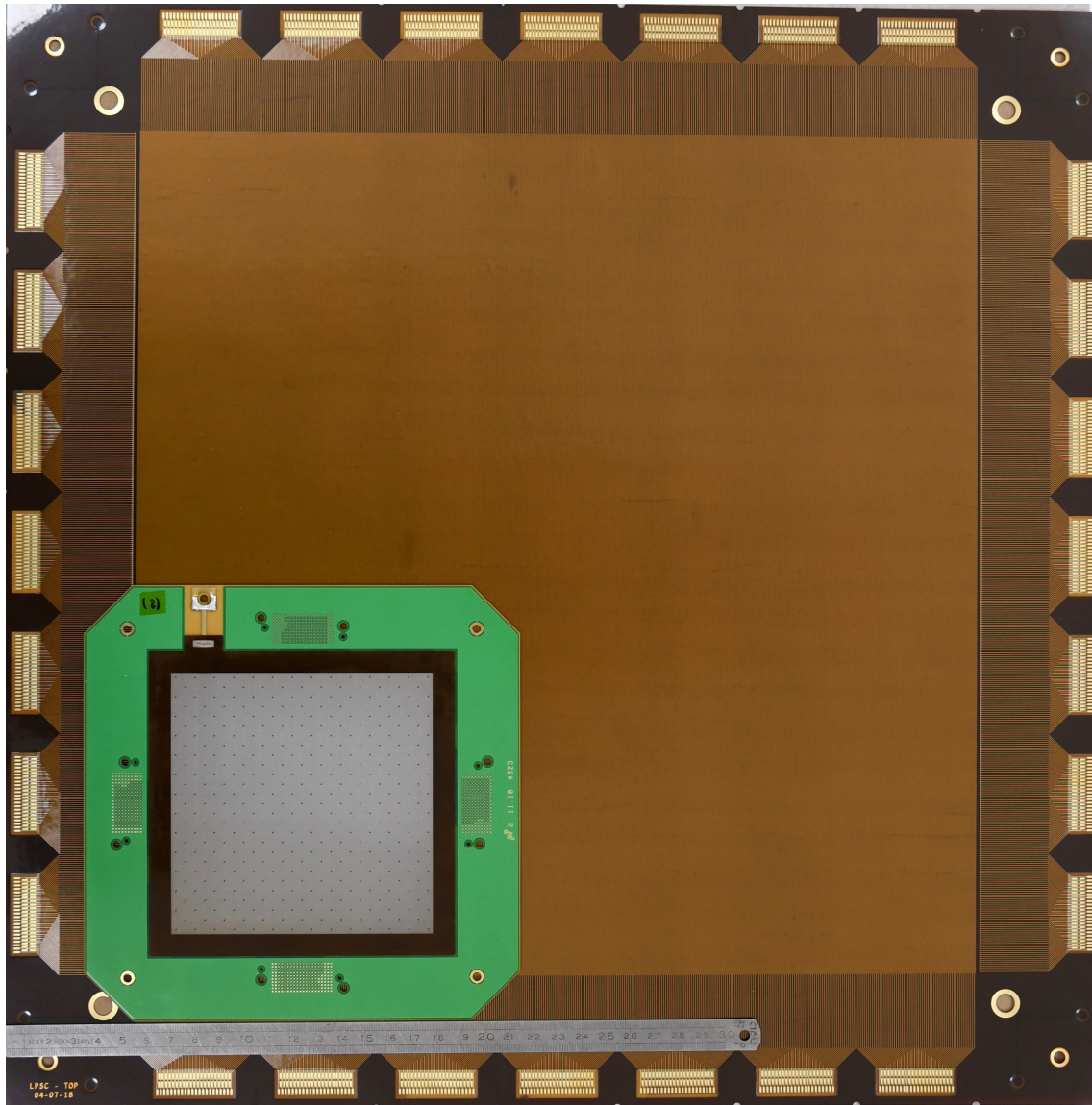


Left: Top view of the new detector design using kapton and plexiglass instead of PCB.

Right: Bottom view, showing the ASICs distribution to minimize the length of the connections.

**The ASIC-MIMAC technology should be upgraded (0.35 to 0.13 μm) :
6 month study for the “electronics” team of the LPSC**

The new 35 cm “new technology” MIMAC detector compared to the old one



Les réalisations techniques envisagées sont :

1) Nouveau module bi-chambre bas-bruit pour les détecteurs 35 cm (voir tableau 1)

Comprenant :

Nouveau détecteur bas-bruit (35 cm x 35 cm) avec la technologie DLC (Diamond Layer Coating) (étude et fabrication en interne)

Nouvelle interface acquisition-détecteur (étude et fabrication en interne)

Nouvelle carte d'acquisition 1792 voies (28 ASICs)

Intégration de nouveaux connecteurs.

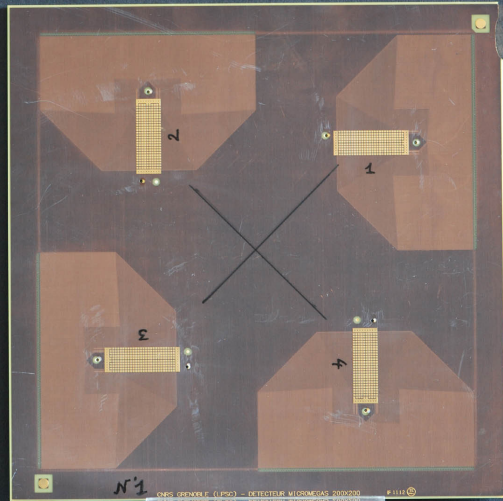
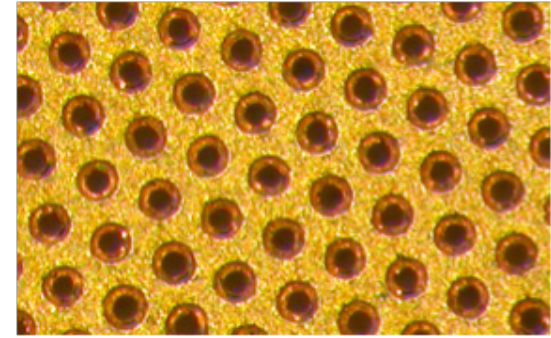
Nouvelle Cage du Champ (en plexiglas et kapton cuivré)

(étude et fabrication en interne)

Filtrage du Rn (en collaboration avec le CPPM-Marseille)

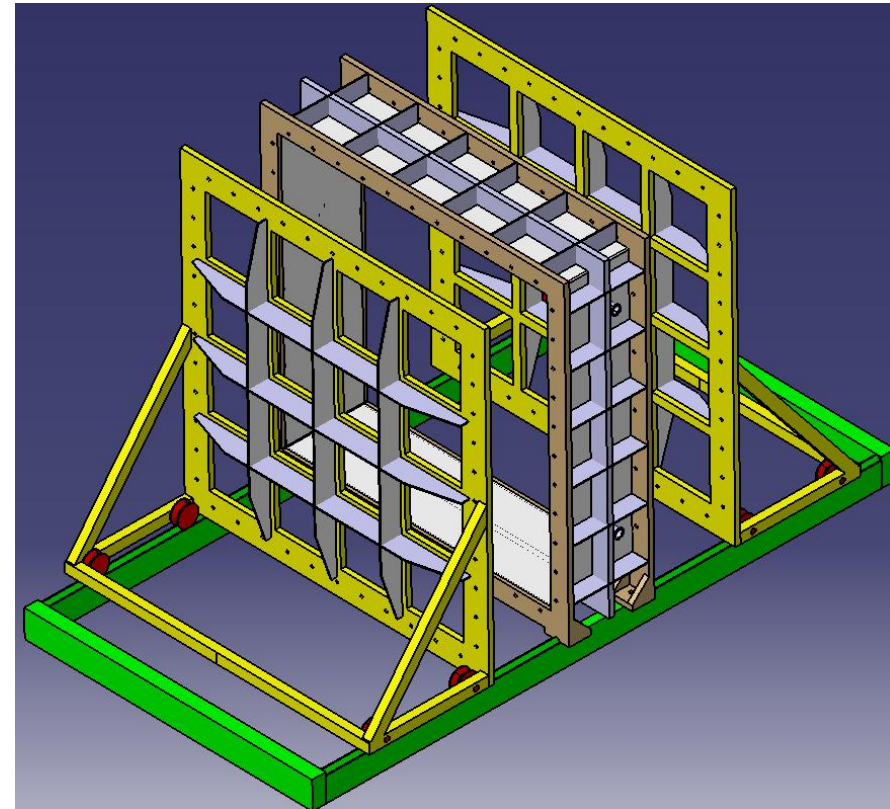
MIMAC – $1\text{m}^3 = 16$ bi-chamber modules ($2 \times 35 \times 35 \times 26 \text{ cm}^3$)

- i) New technology anode $35\text{cm} \times 35\text{cm}$
- ii) Stretched thin ($12 \text{ }\mu\text{m}$) grid at $512\text{ }\mu\text{m}$.
- iii) New electronic board (1792 channels)
- iv) Only one big chamber



New $20\text{cm} \times 20\text{cm}$ pixellized anode
(1024 channels)

CS-IN2P3, Paris, 25 octobre 2018



Analyse “SWOT”

Forces :

MIMAC a bénéficié d'un développement « maison (IN2P3)»

La capacité de mesurer de très faibles énergies en ionisation et de reconstruire des traces de longueur sub-millimétrique.

La possibilité de développer des détecteurs MPGD bas bruit.

Nous maîtrisons le savoir-faire : détecteur, électronique et mesure du facteur de quenching en ionisation.

L'accès et la proximité au LSM (Laboratoire Souterrain de Modane).

Faiblesses :

Les équipes de physiciens du LPSC et du CPPM sont sous-dimensionnées.

Une autre équipe intéressée à participer serait souhaitable !

Opportunités :

Notre stratégie de détection et technologie MIMAC sont pour l'instant celles qui permettent le seuil le plus bas avec la mesure de traces en 3D.

Valorisation de la technologie dans le domaine de la mesure neutronique.

Menaces :

Les particules de matière sombre pourraient être d'une masse inférieure à 500 MeV/c².
Les particules de matière sombre pourraient être d'une masse inférieure à 500 GeV/c².

Conclusions

- A new directional detector of nuclear recoils **at low energies ($E > 100$ eV)** has been developed giving a lot of flexibility on targets, pressure, energy range...
- New observables are available to discriminate electrons from nuclear recoils to improve the DM search for and to cope with « non zero » background.
- Angular resolution and directional studies of 3D tracks have been performed experimentally with COMIMAC showing large discrepancies wrt simulations!
- **The 1 m³ will be the validation of a new generation of a large DM high definition detector including directionality (a needed signature for DM discovery)**
- **The « large volume » needed for Dark Matter detection could be possible to get it by an array in different underground laboratories !**

We need to help the Grenoble and Marseille teams to keep the know-how and succeed